

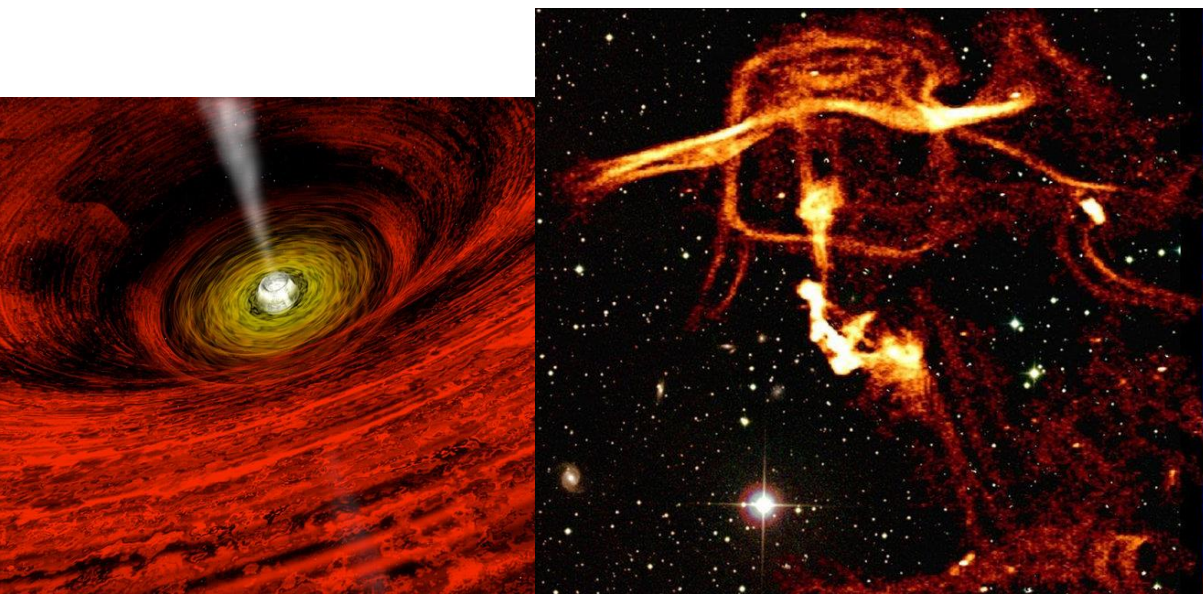
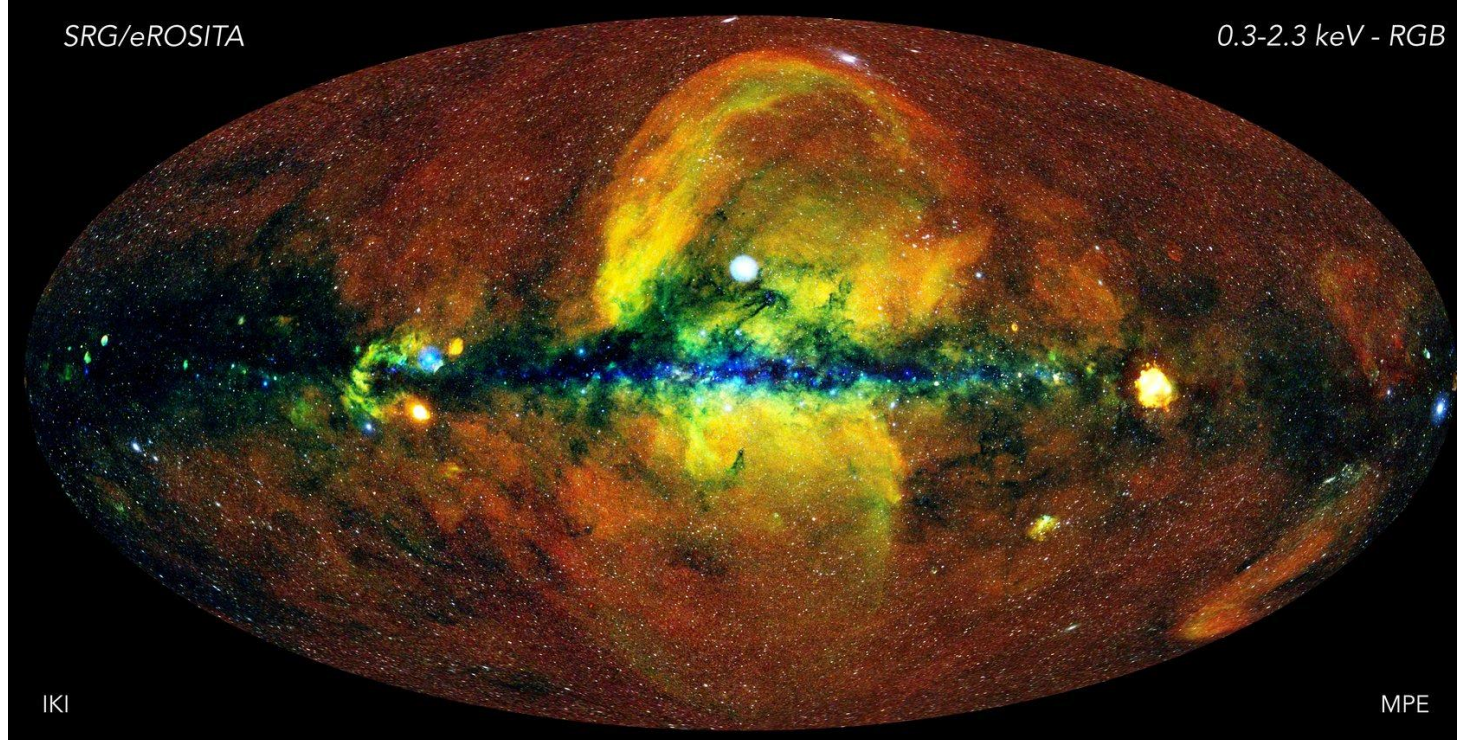


Copernicus 550 Meeting Torun, Poland

February 21, 2023

Map of the entire sky in X-rays

31 months of the work in space

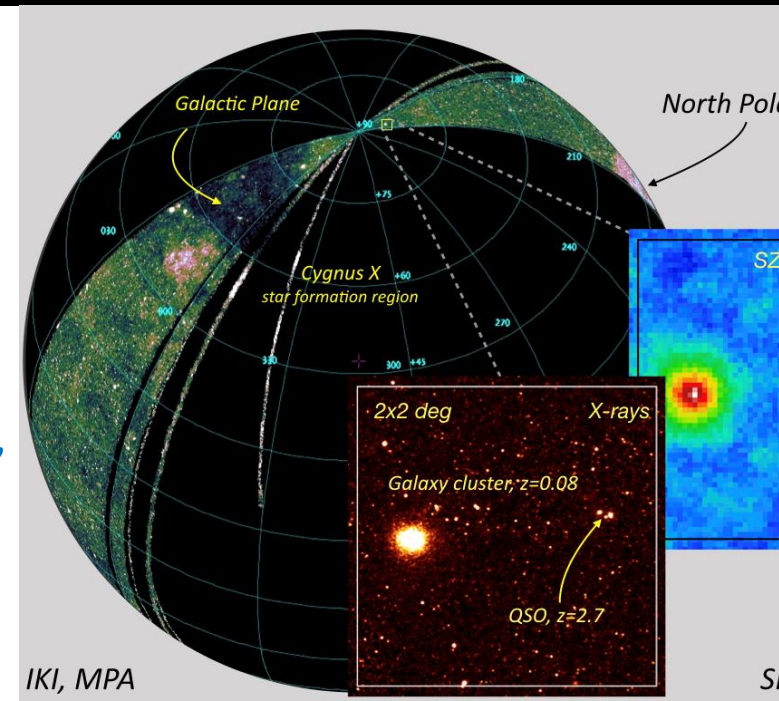


Rashid Sunyaev

Max-Planck Institute
for Astrophysics

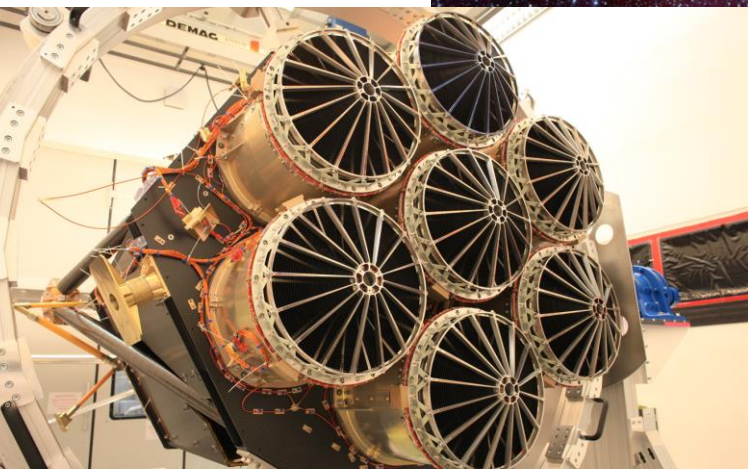
Institute for Advances Study,
Princeton

Space Research Institute,
Moscow



Spectrum-RG (SRG, 2700 kg) Orbital Observatory

Located near L2
Point of the
Sun-Earth system
at the distance 1.5
Mln km from the
Earth not very far
from JWST (NASA)
and Gaia (ESA)



eRosita, 808 kg
1.9 m * 3.5 m

eRosita (MPE)
has seven parallel
grazing incidence
mirror systems and
seven X-Ray CCDs.
Field of view is 1
degree, angular
resolution 25"

0.3 – 9 KeV
Pixel - 9.6" * 9.6"

ART-XC hard X-Ray Telescope (350 kg), IKI (Moscow), All-Russia
Nuclear Center (Sarov) and Marshall Space Flight Center (NASA;
X-Ray grazing incidence mirrors)

4 – 30 KeV

NAVIGATOR platform

Lavochkin industry
(near Moscow)
Launch 13.07.2019

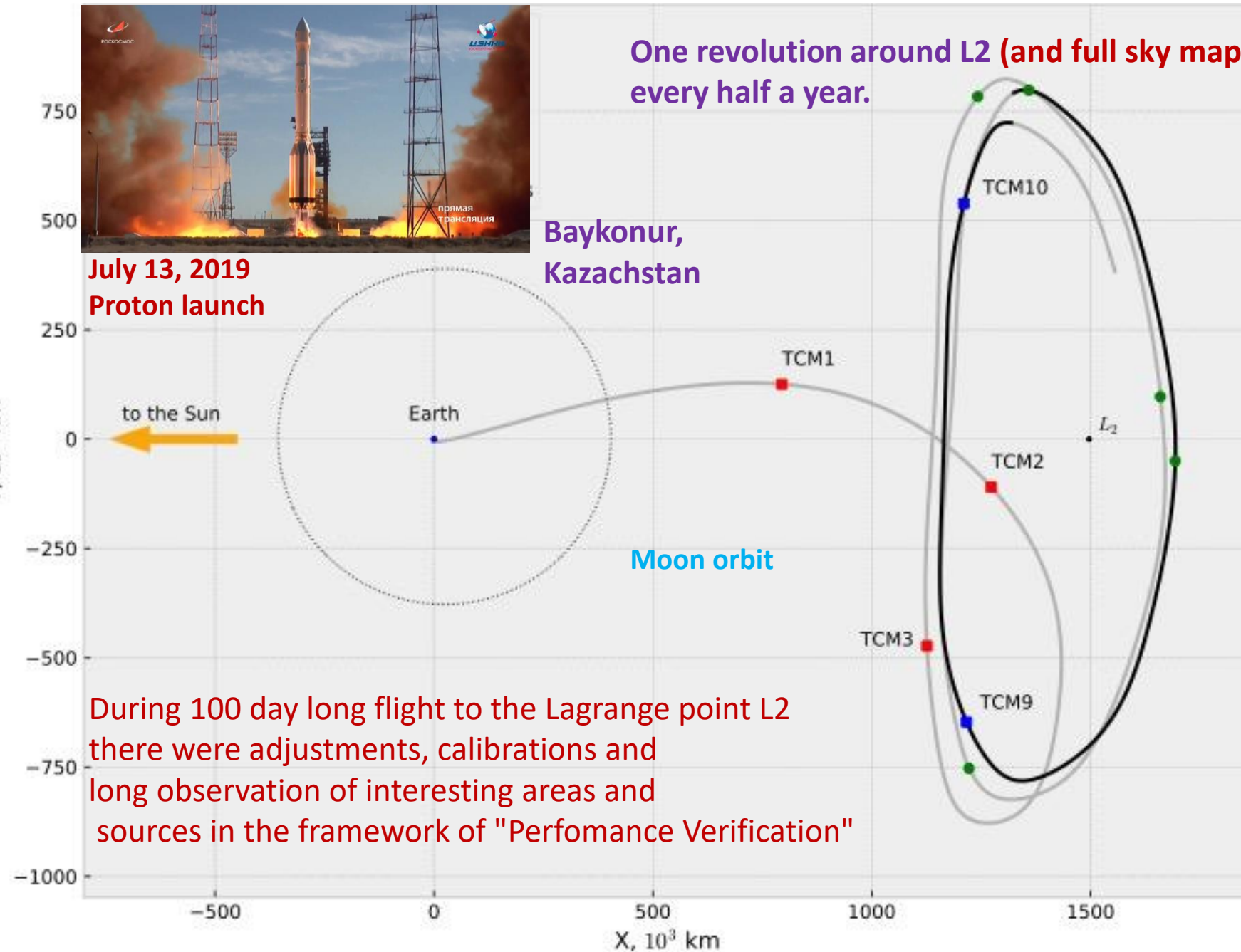
Projection of the SRG trajectory onto the ecliptic plane



July 13, 2019
Proton launch

Baykonur,
Kazakhstan

**One revolution around L₂ (and full sky map)
every half a year.**

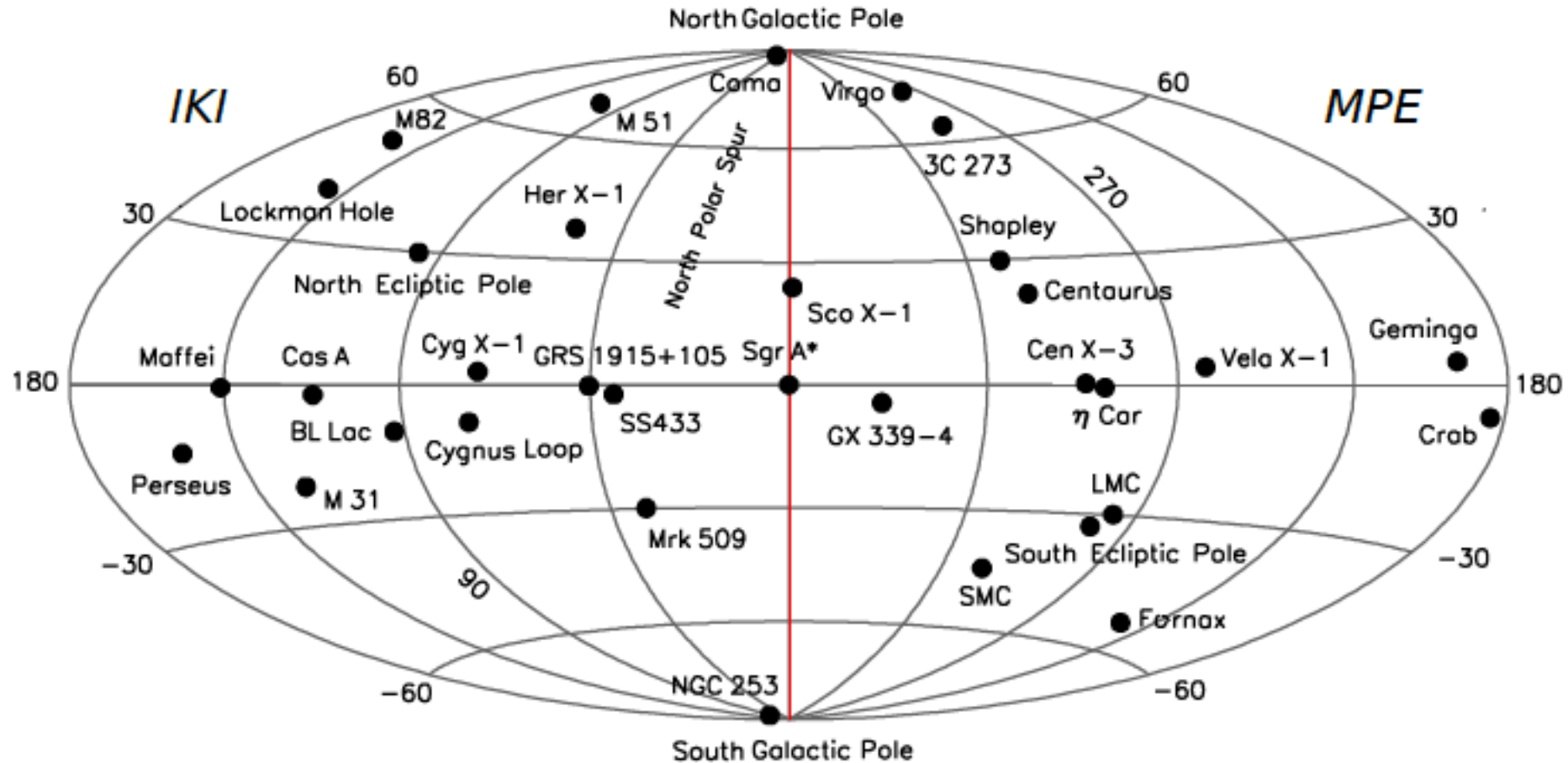


**During 100 day long flight to the Lagrange point L₂
there were adjustments, calibrations and
long observation of interesting areas and
sources in the framework of "Performance Verification"**

2007: ROSKOSMOS and DLR (German Air and Space Agency) made the statement and included it into the Memorandum of Understanding:

for the data processing and publications of SRG/eRosita results **on one half of the sky will be responsible Russian scientists. German scientists will be responsible for results from another hemisphere.**

Galactic coordinates



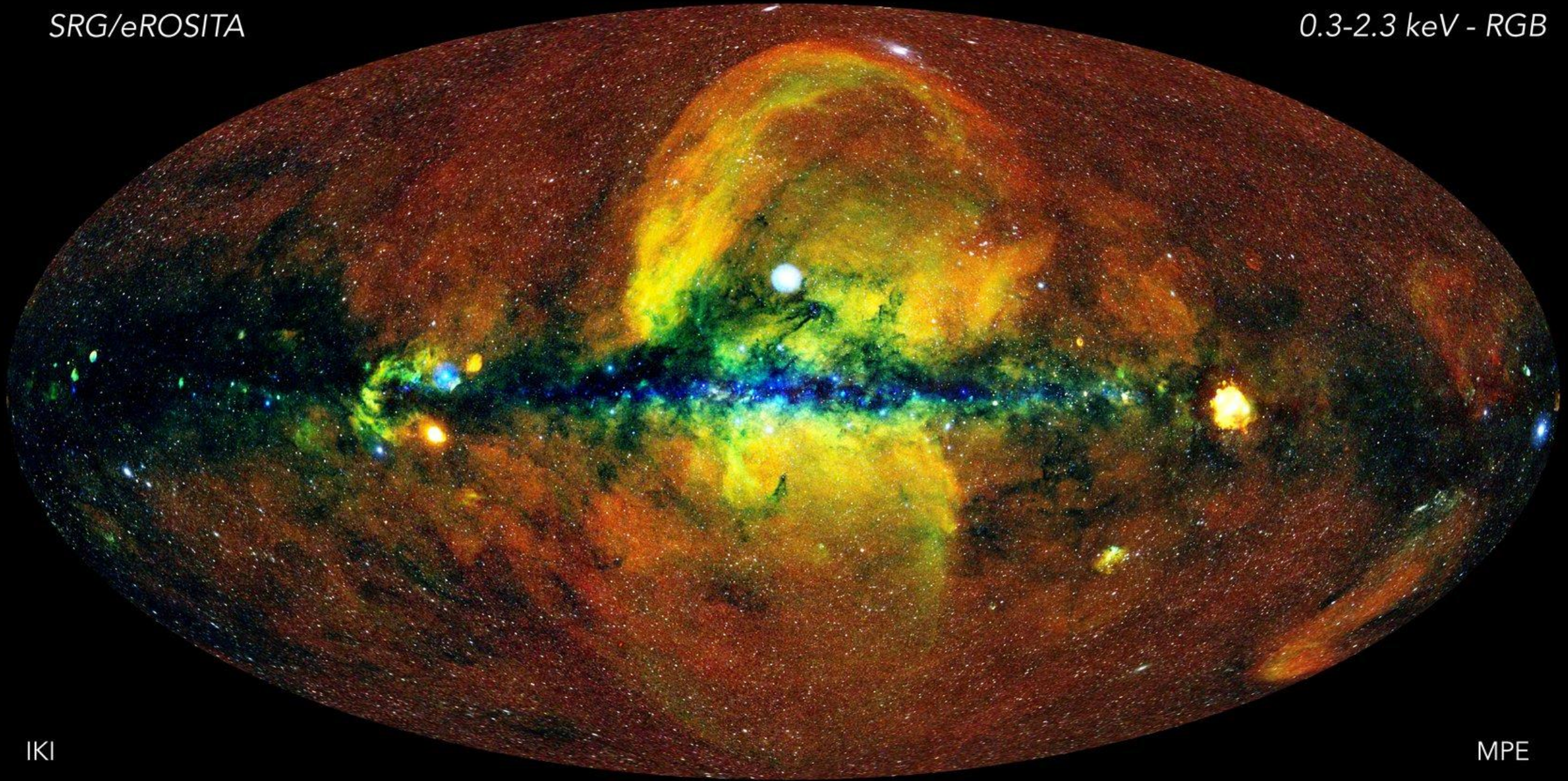
IKI – Space Research Institute
of the Russian Academy of Sciences

MPE – Max-Planck Institute
fuer Extraterrestrische Physik

X-Ray Map of the entire Sky

SRG/eROSITA

0.3-2.3 keV - RGB



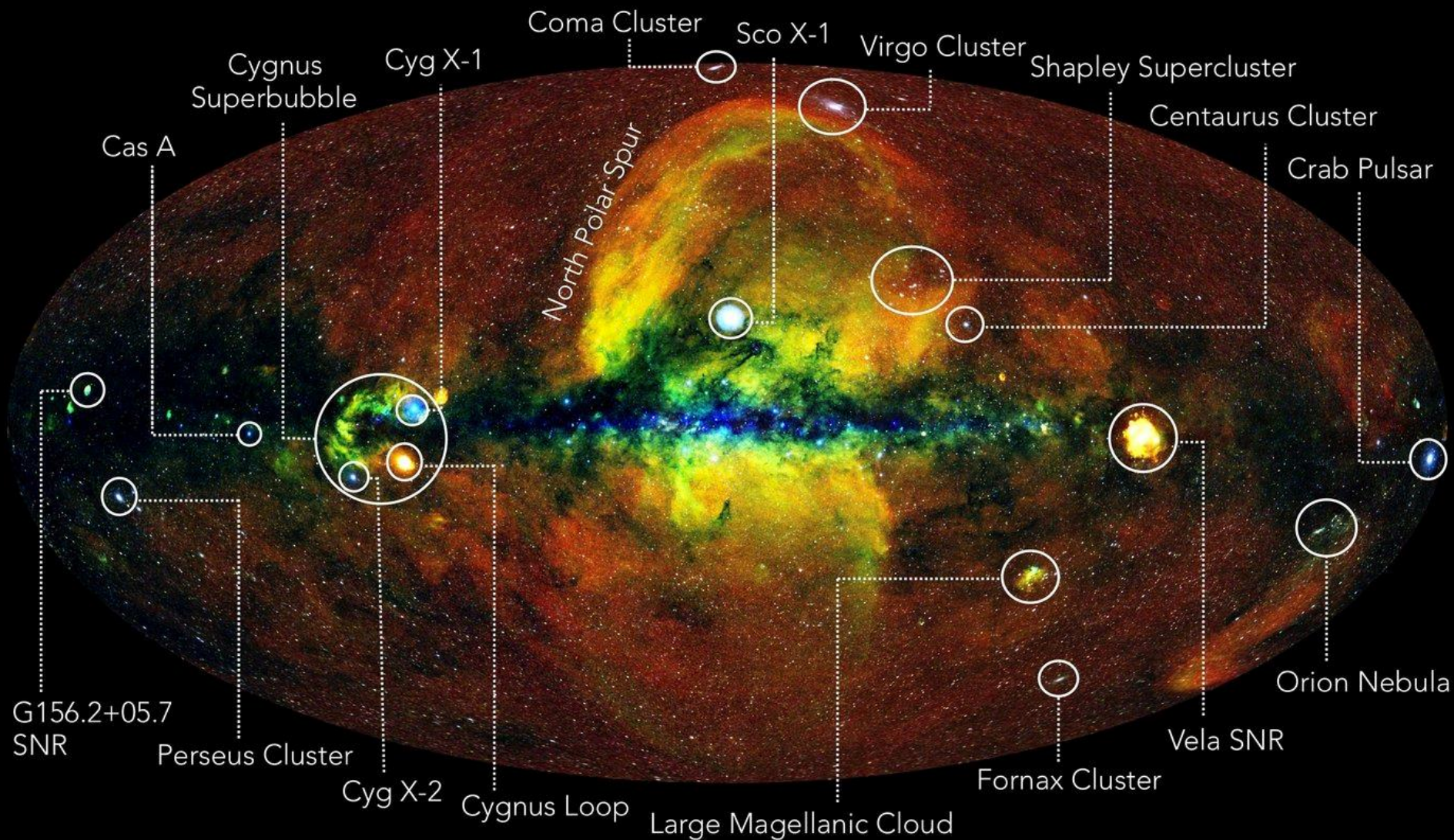
IKI

MPE

Churazov, Gilfanov

Brunner, Sanders

Navigating the eROSITA X-ray sky



IKI

Churazov, Gilfanov, Sunyaev

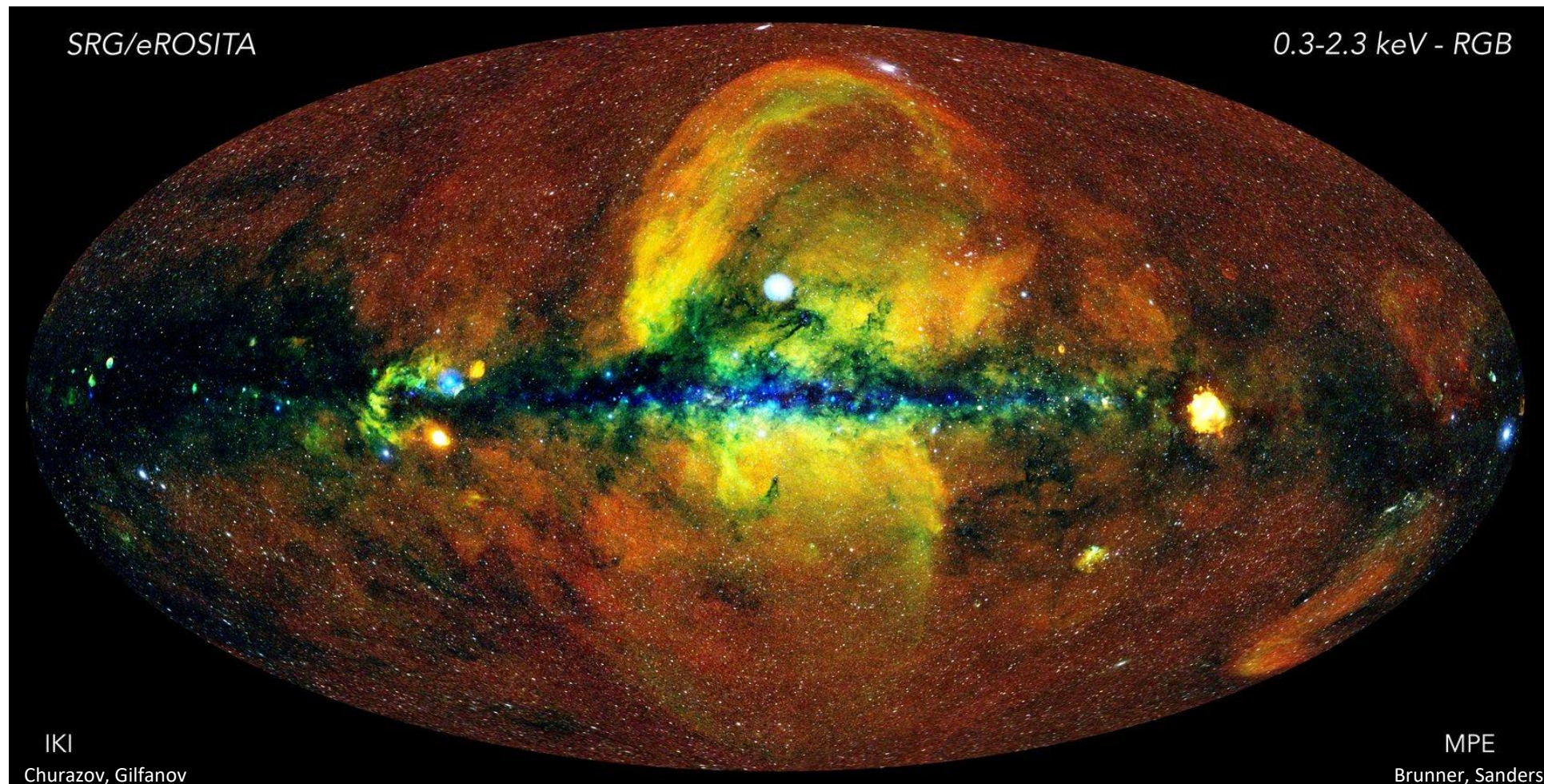
SRG/eROSITA 0.3-2.3 keV - RGB Map

MPE

Brunner, Merloni, Sanders

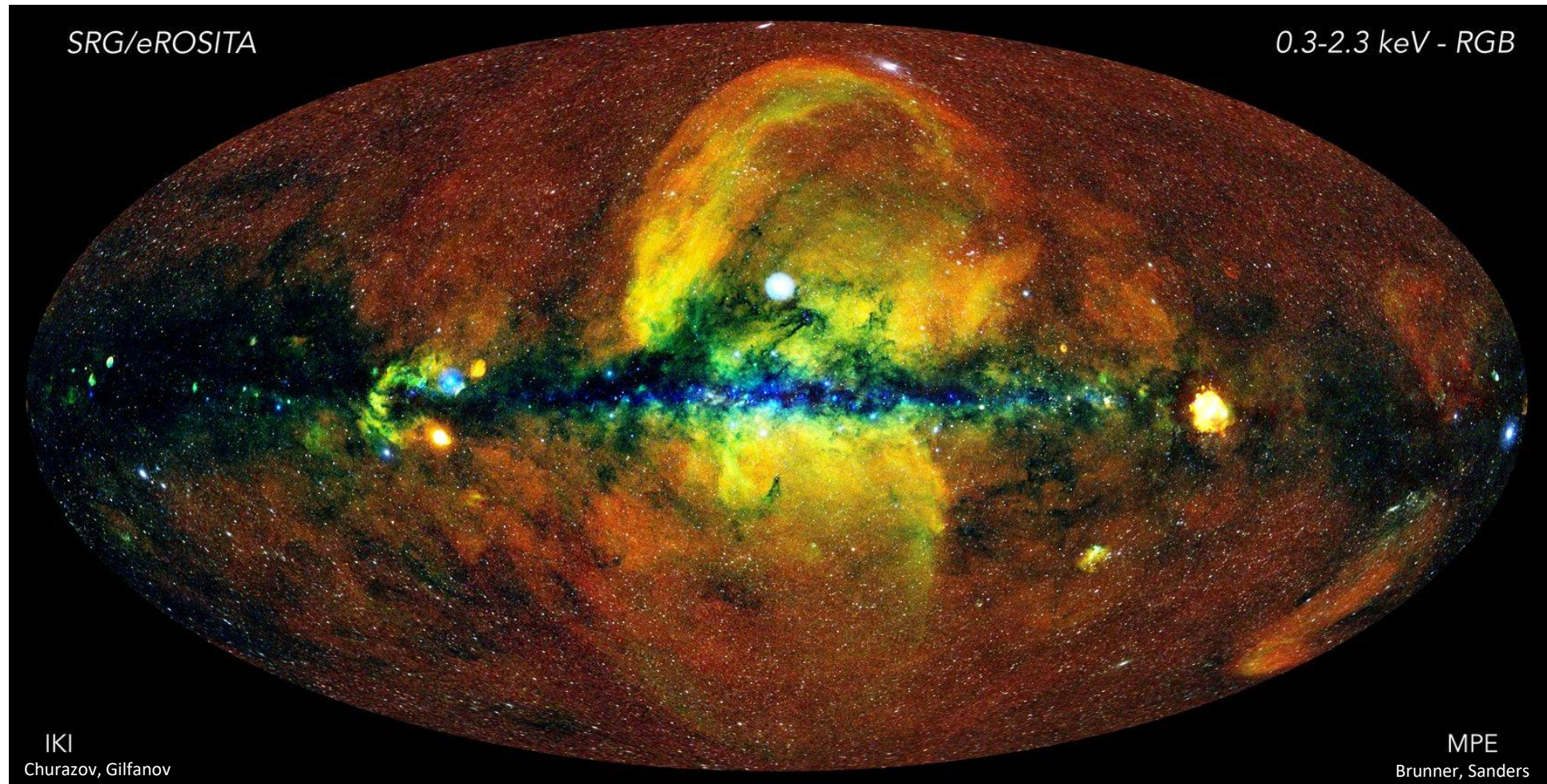
First all-sky map of Rosita:

Million of X-Ray sources, Milky Way
and traces of huge explosions at the Center of our Galaxy



This is a three colour RGB map. All **400 millions of photons** with energies from 300 eV up to 2.3 KeV detected during half a year of continuous scanning are presented on this map. Red colour corresponds to photons with energies 0.3–0.6 KeV, green -- 0.6–1 KeV and blue – to 1–2.3 KeV. (Today after **26 months** of scanning the sky we have **more than one and a half billion of X-Ray photons**).

The first SRG all-sky survey allowed to construct a map containing almost 8 times more X-Ray sources than the former world-best map of the ROSAT satellite, obtained in 1990.



Three quarters of a Million objects on this map are distant quasars and active galactic nuclei powered by accretion of matter onto supermassive black holes residing in their centers. They are far beyond the Milky Way at distances of hundreds of millions and billions of light years from us. We see also 20 000 extended objects (mainly clusters and rich groups of galaxies with hot gas) and more than 200 000 galactic stars with active coronae. 400 Millions photons



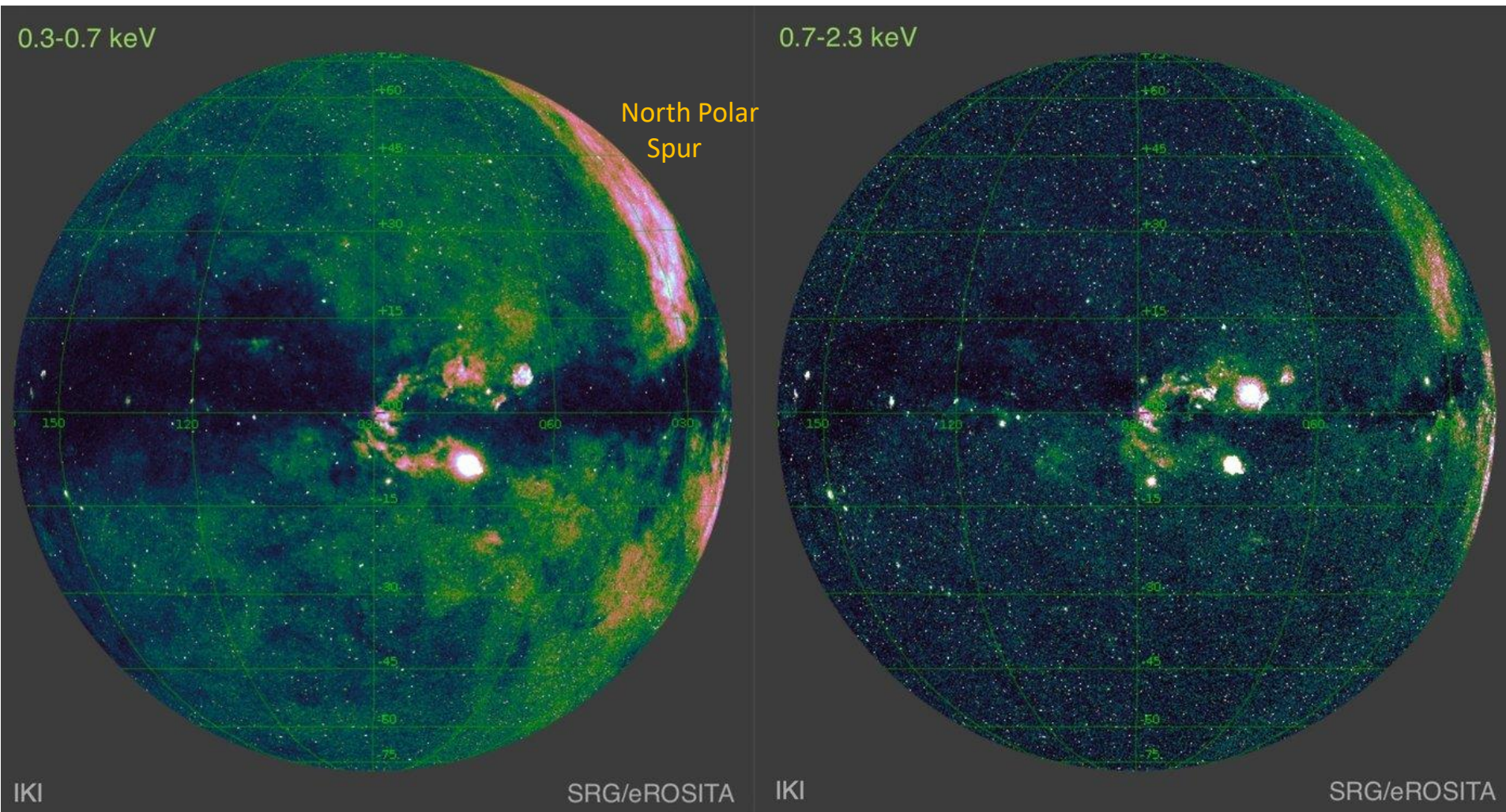
Marat Gilfanov

The SRG/eRosita maps of the hemisphere analyzed by the Russian eROSITA consortium

Left side: hot gas emission above the Plane of the Galaxy (halo, local Bubble?, boundary of Heliosphere?).

Strong absorption of X-Rays by the cold gas in the Plane.

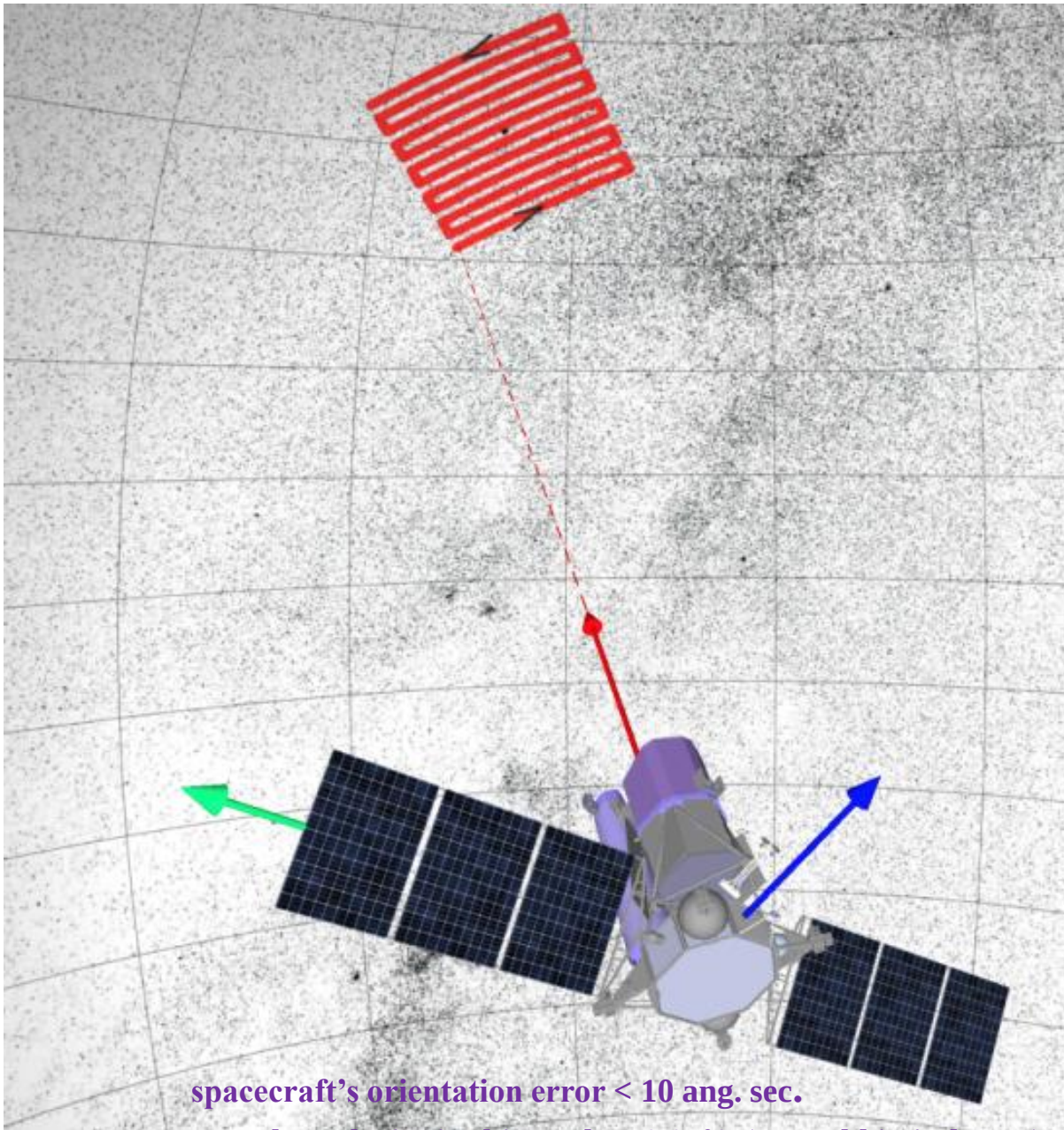
Right side: much more X-Ray sources on cosmological distances. Hot gas has temperature less than few million K. Weaker absorption near the Plane of the Galaxy (just cross section).



scan of selected regions up to $12.5^\circ \times 12.5^\circ$ in size

Lockman Hole

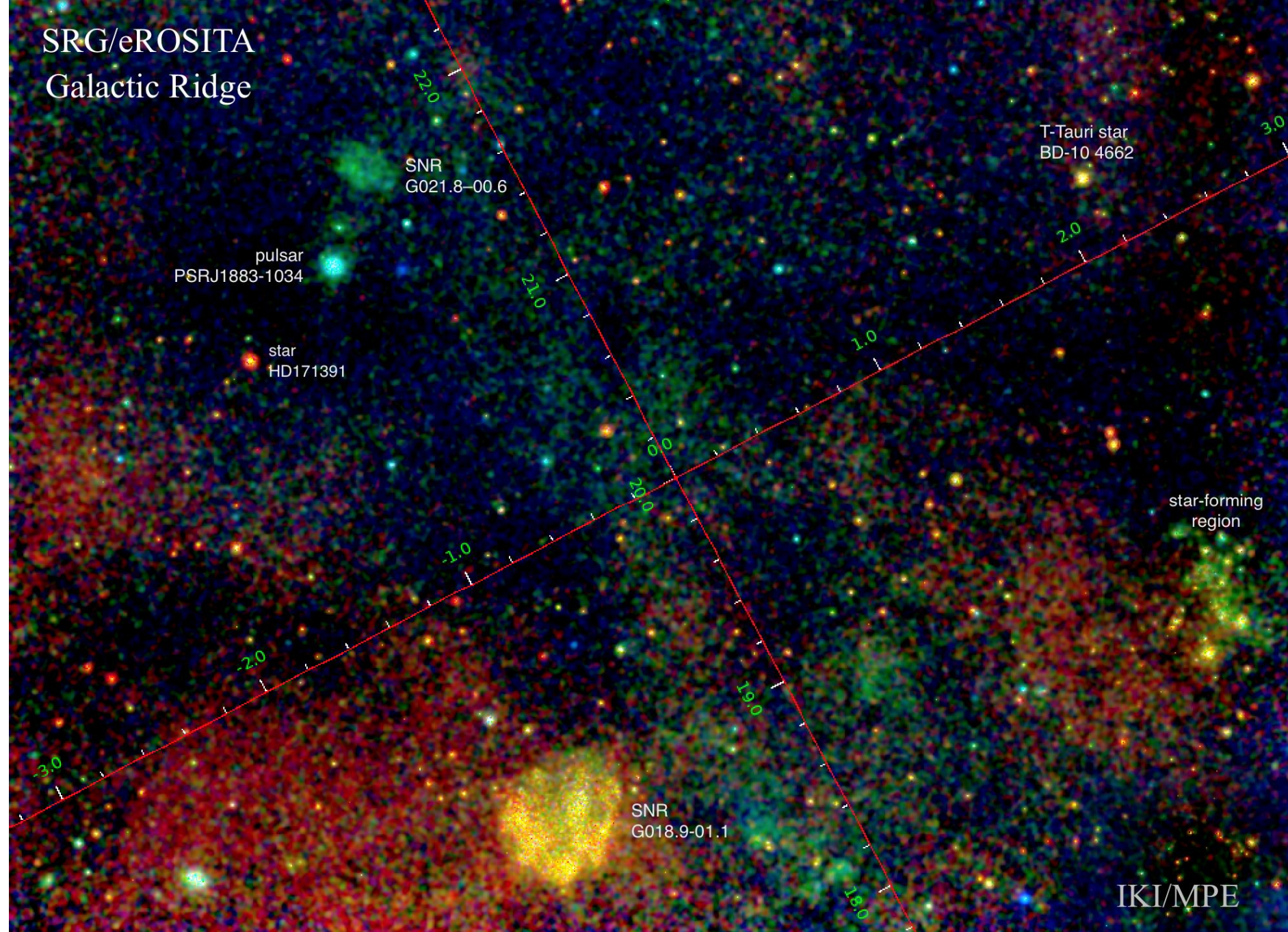
СРГ/«РОЗИТА»
“дыра Локмана”



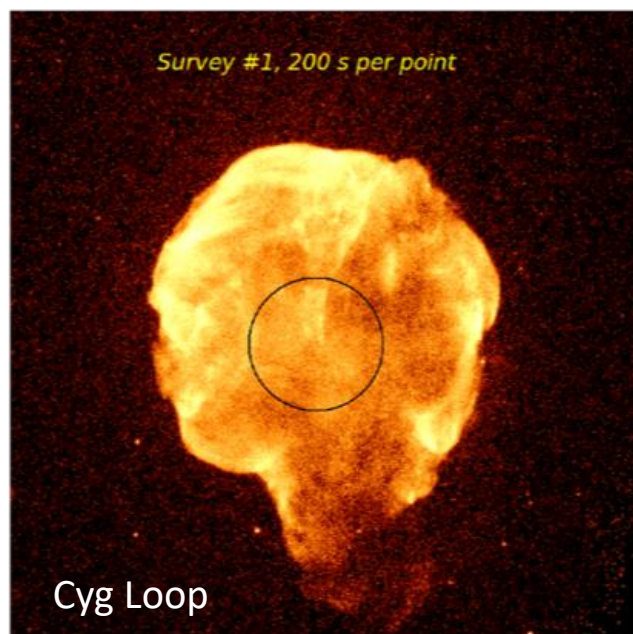
spacecraft's orientation error < 10 ang. sec.
angular velocity in this mode $< 0.04^\circ/\text{s}$ (Lavochkin industry)



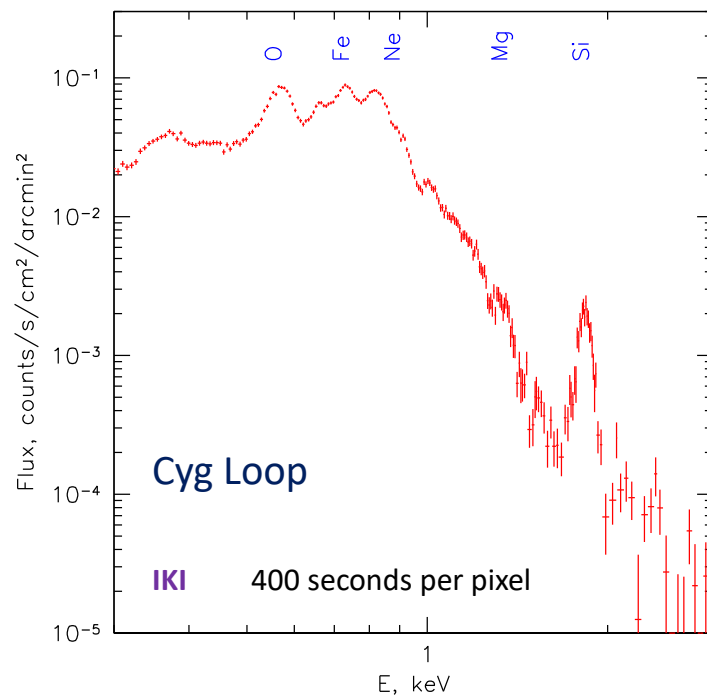
20 sq deg: 6000 X-Ray sources, 200 extended,
185 confirmed clusters of galaxies (spectroscopic
 z for 150); Rodion Burenin, Marat Gilfanov, Rodion
Burenin, Sergej Bykov et al



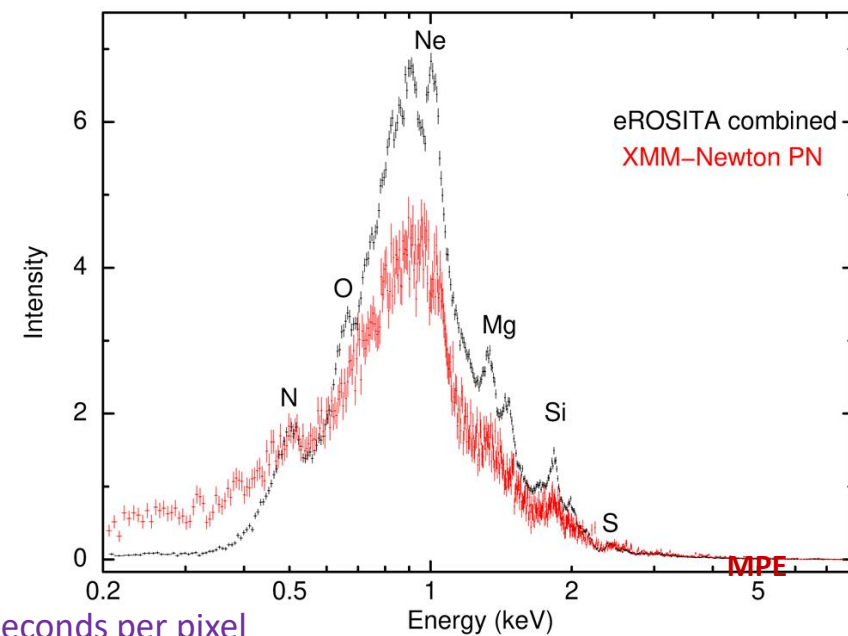
25 sq deg in 20 degrees from the Galactic Center (Marat Gilfanov et al)



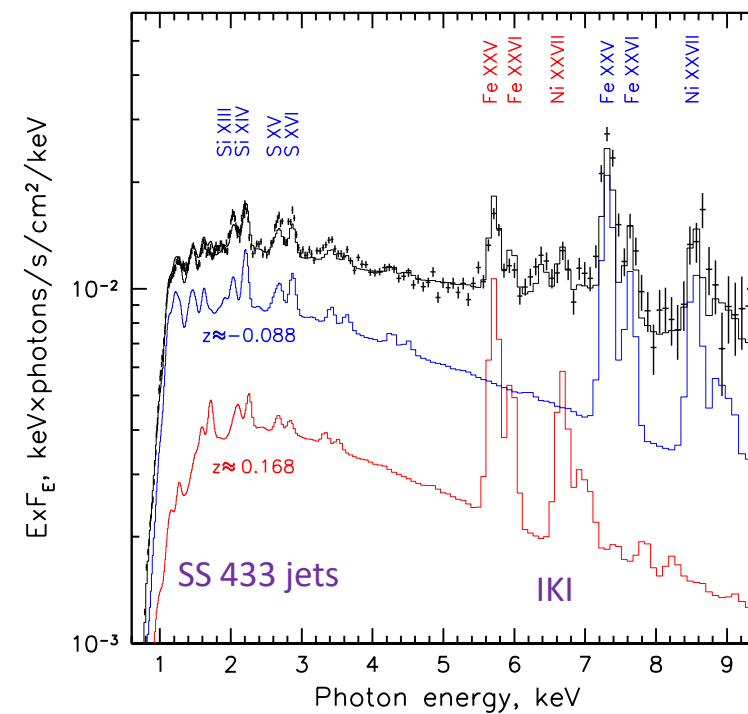
Spectroscopy capabilities of SRG/eRosita



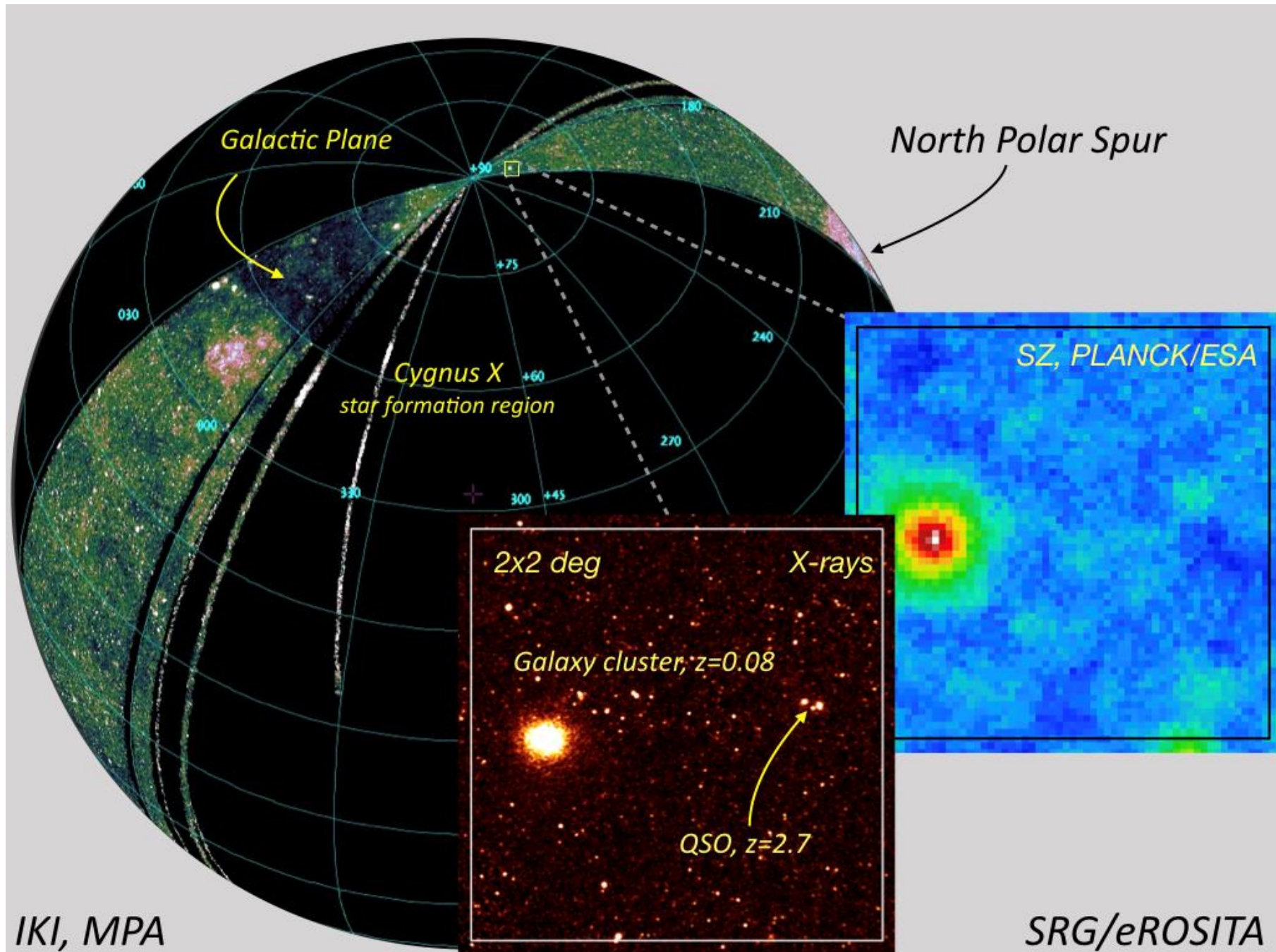
SN 1987A in the LMC



several Kseconds per pixel

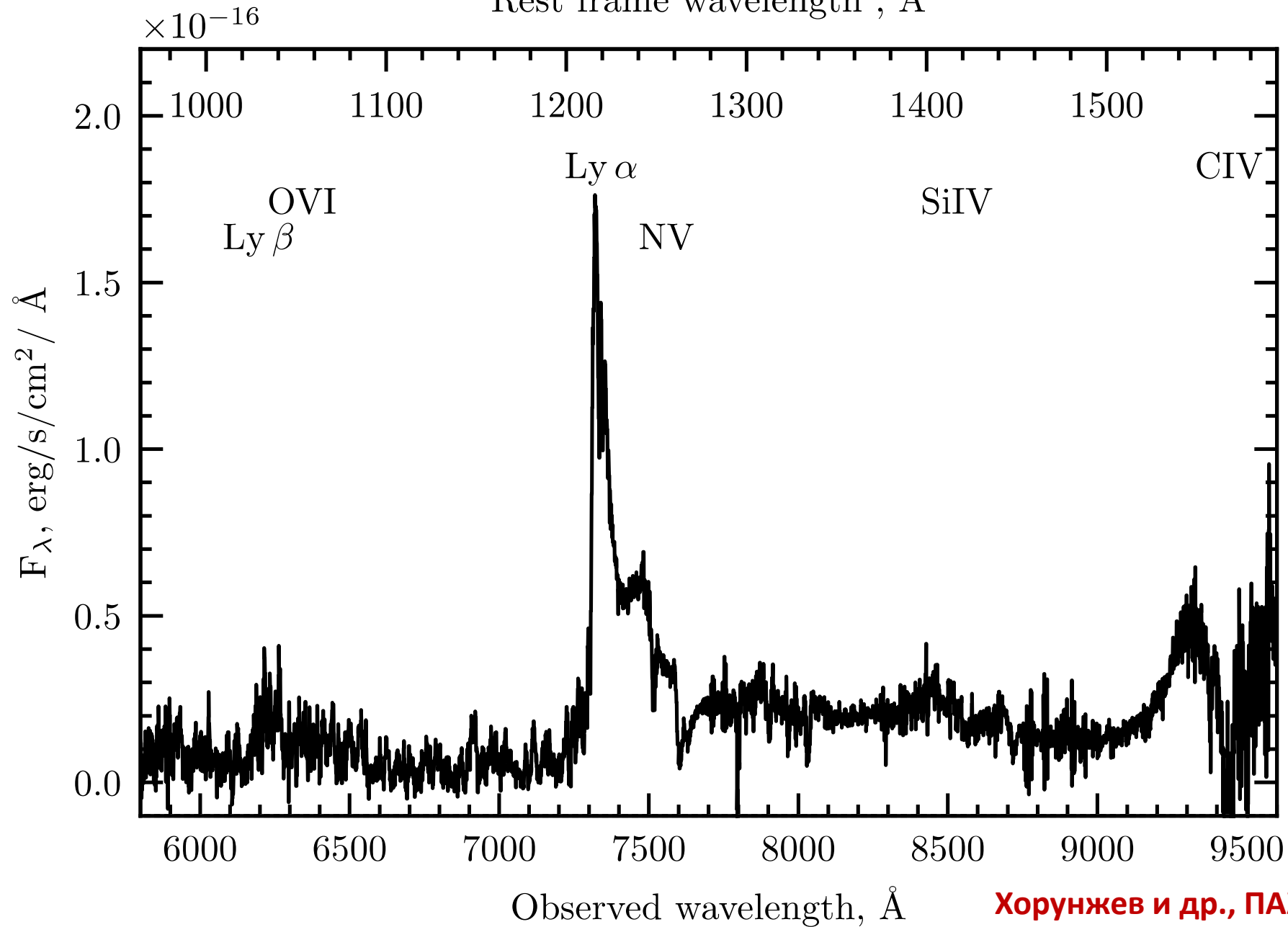


Now whole sky is covered four times and 10 % of the sky is covered during current fifth survey



SRGeJ020142.8-015347 $z_{sp}=5.02$

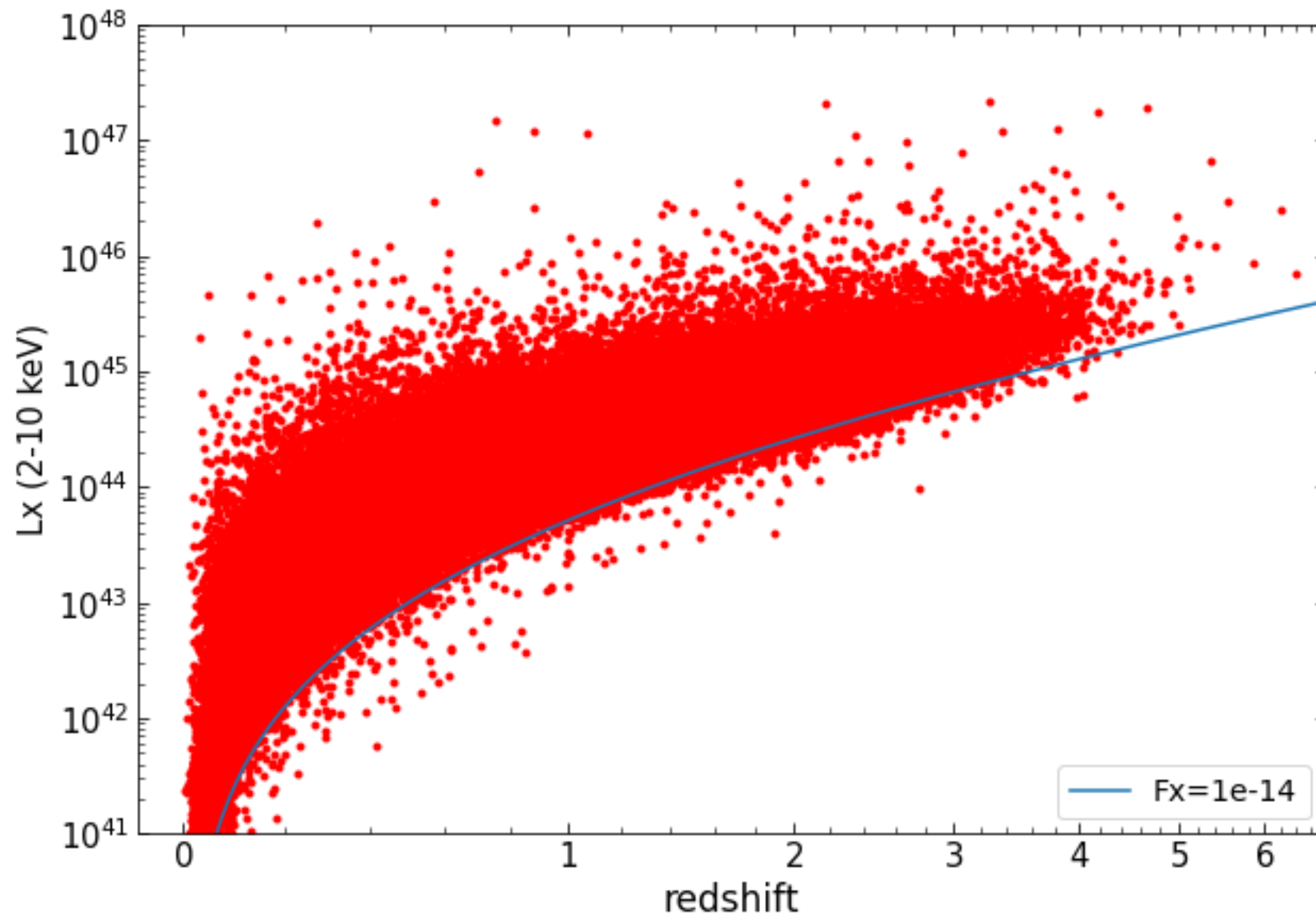
Rest frame wavelength, Å



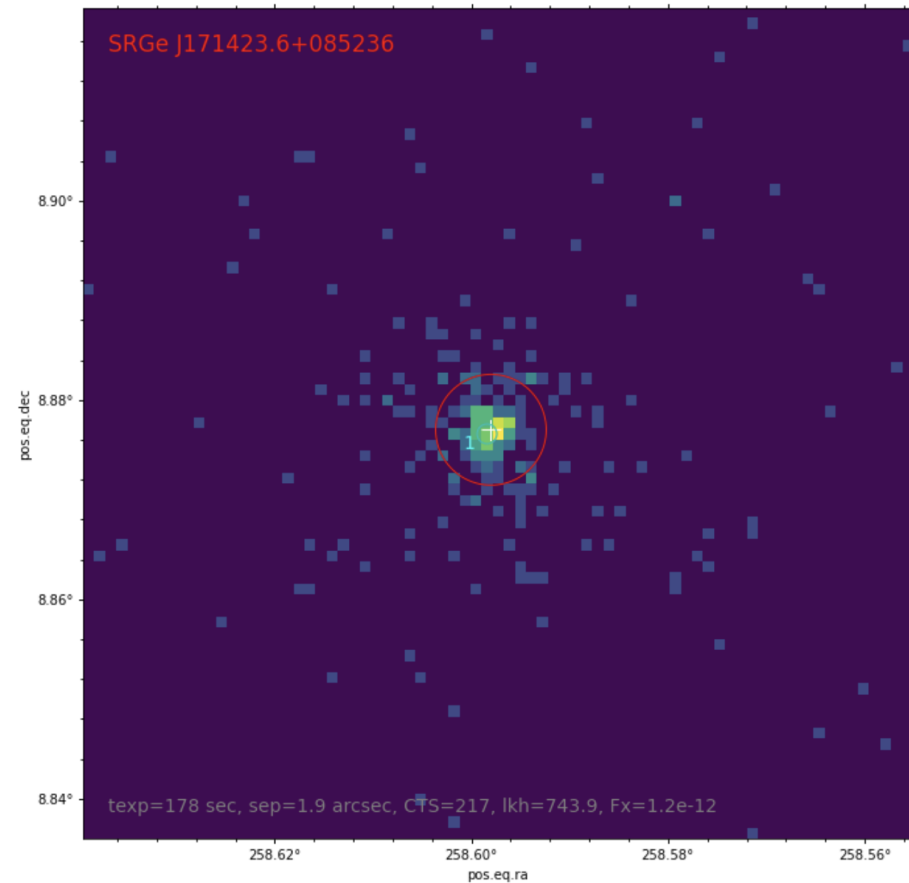
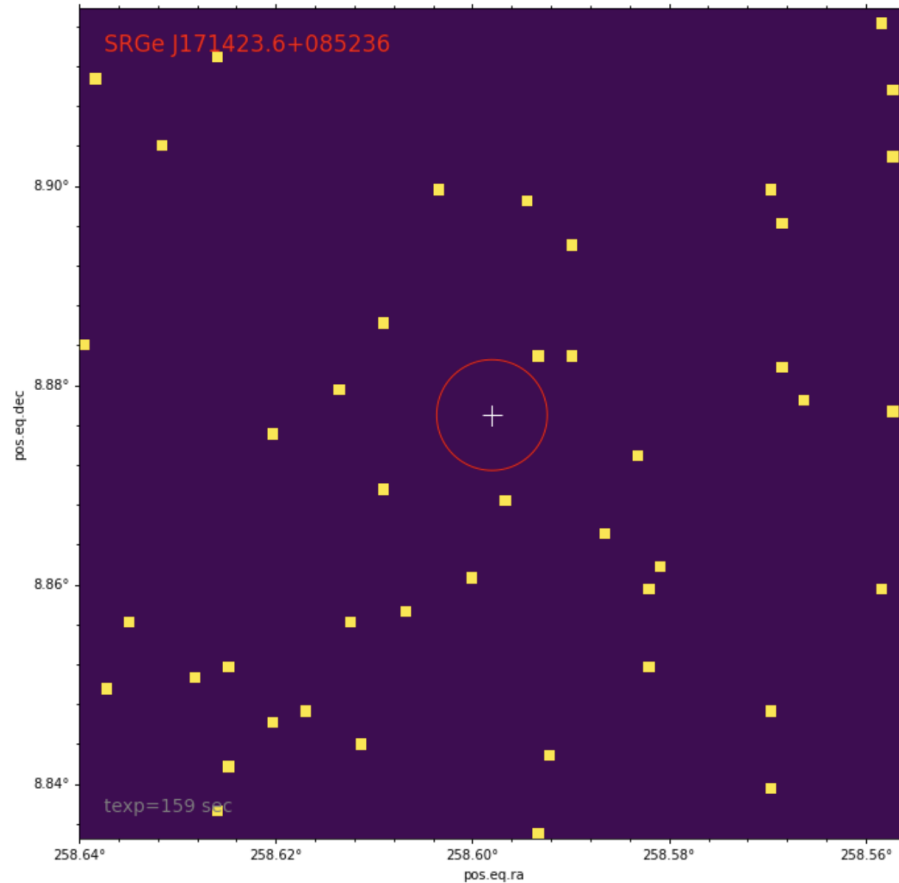
*Павел Медведев открыл
рентгеновский квазар с
красным смещением $z=6.18$*

Хорунжев и др., ПАЖ, 2, 2022

БТА, 6 м



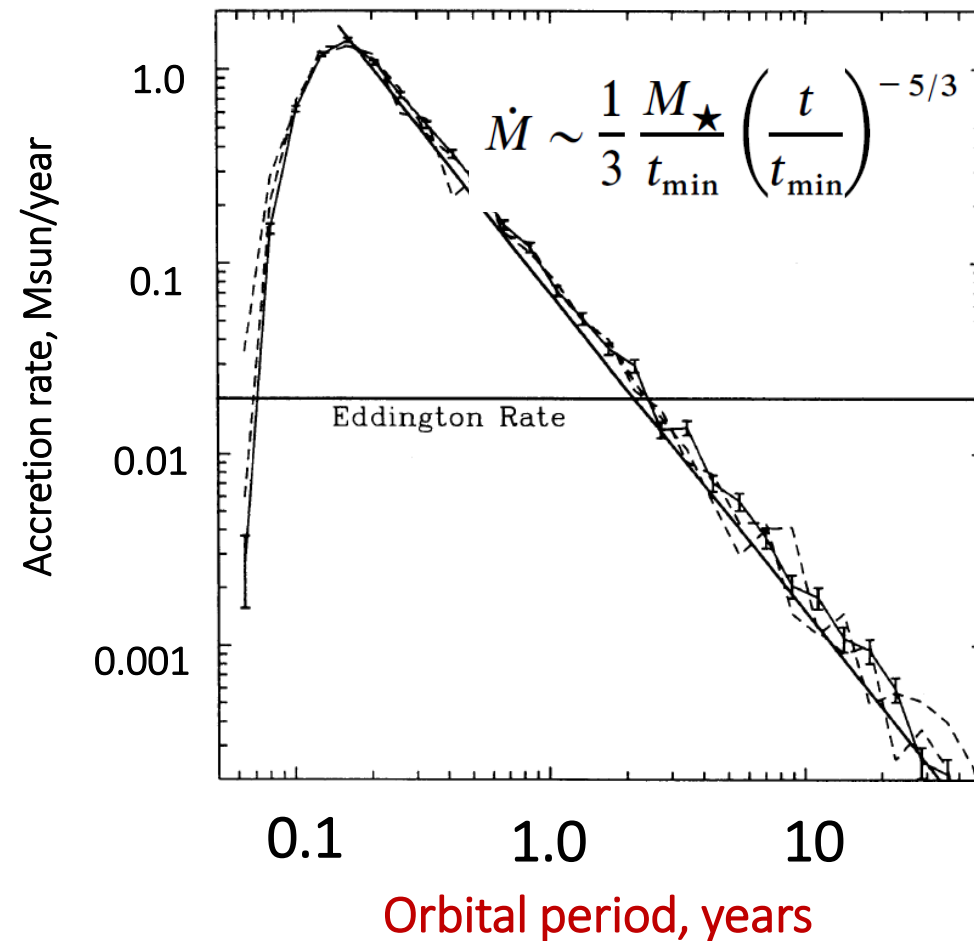
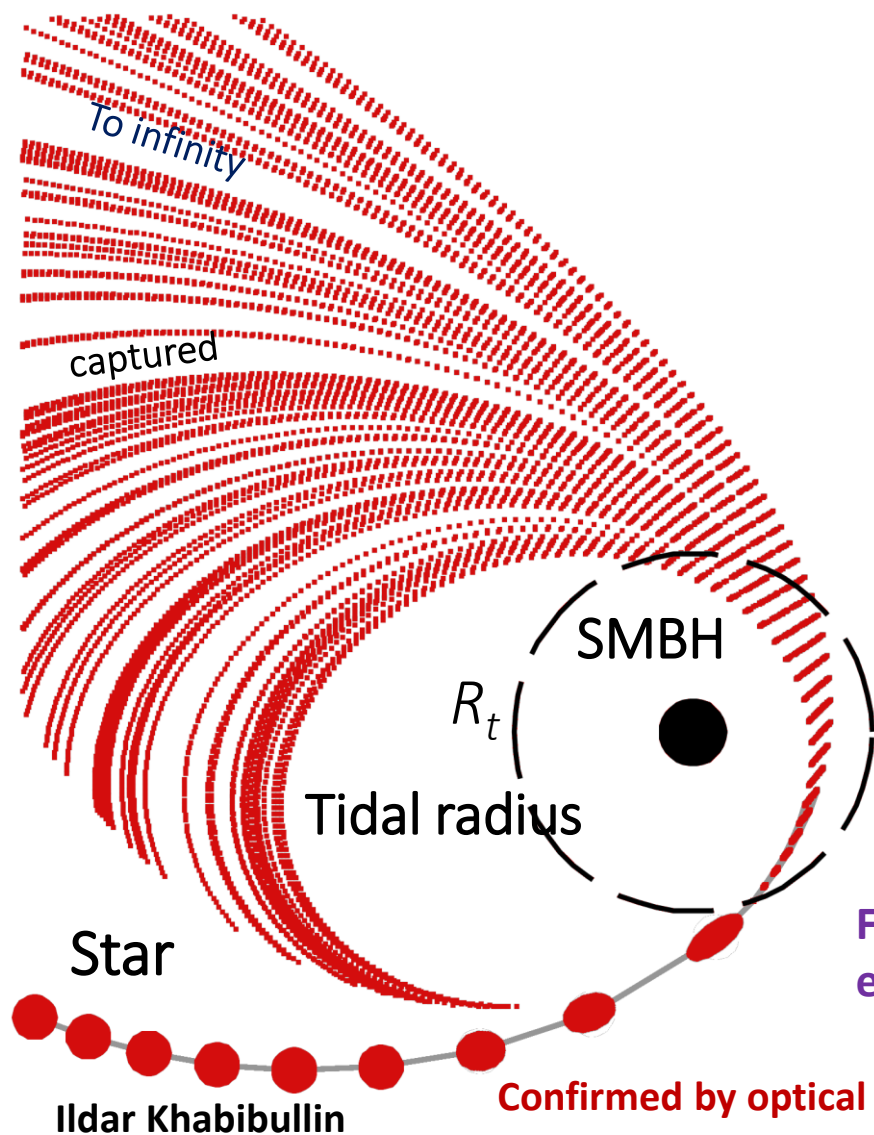
Extragalactic transients



Typically we **detect ~3-5 objects per day changing their flux by >10 times**
(stars, AGN, X-ray binaries in the Milky Way).

On average, we detect **one good TDE candidate per ~10 days**. They are
being followed up by telescopes in Russia, Palomar, Keck, ZTF

Tidal star disruption by a supermassive black hole (TDE)



We have now more than 50 good TDE candidates

First catalog of 13 TDE events, discovered by SRG/
eRosita. S. Sazonov, M. Gilfanov et al, MN RAS, 2021

Confirmed by optical spectroscopy, BTA, RTT150, AZT33IK, Keck, Palomar etc

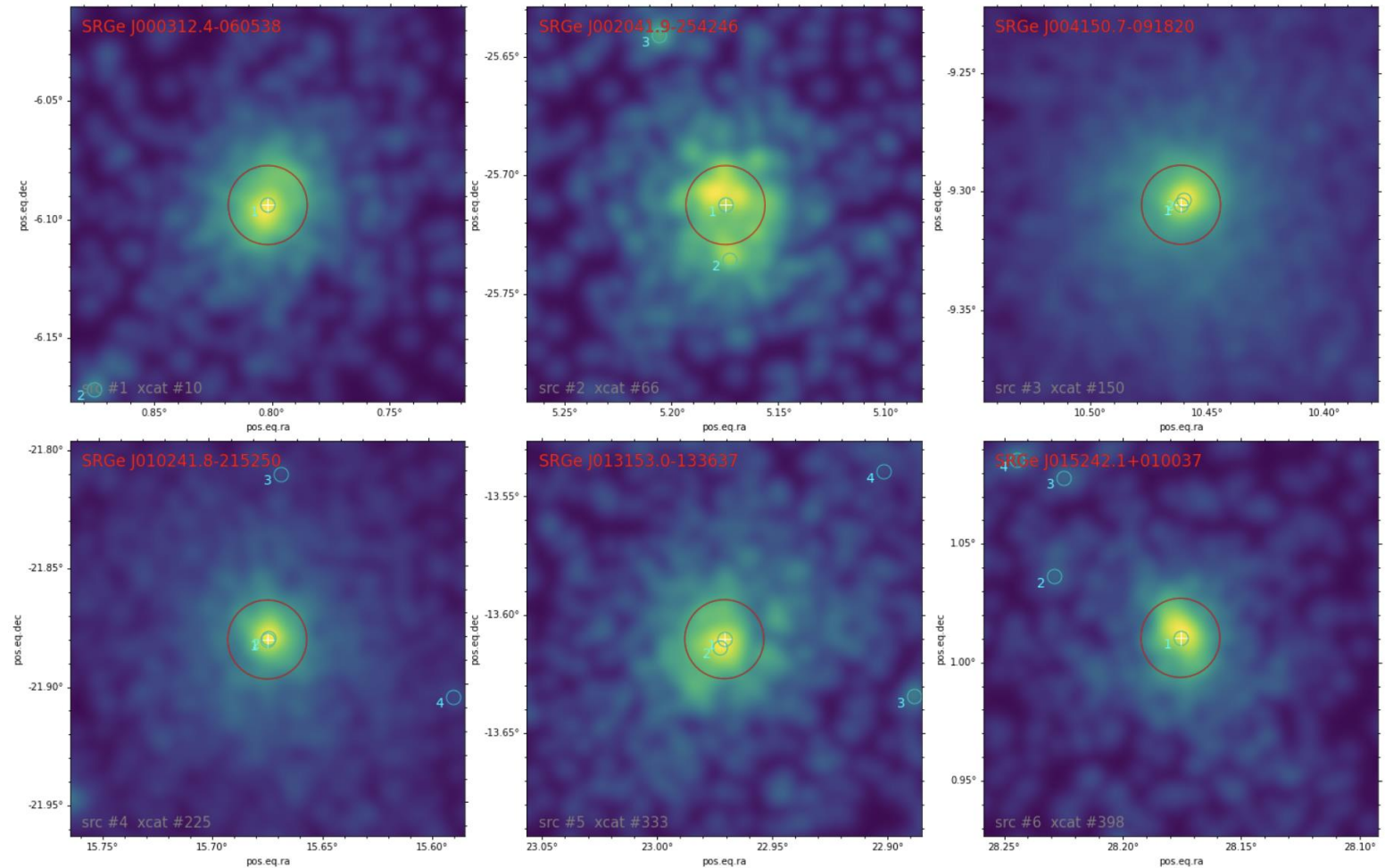
Clusters of galaxies,

discovered by SRG/eRosita
as extended X-ray sources

Half of the sky

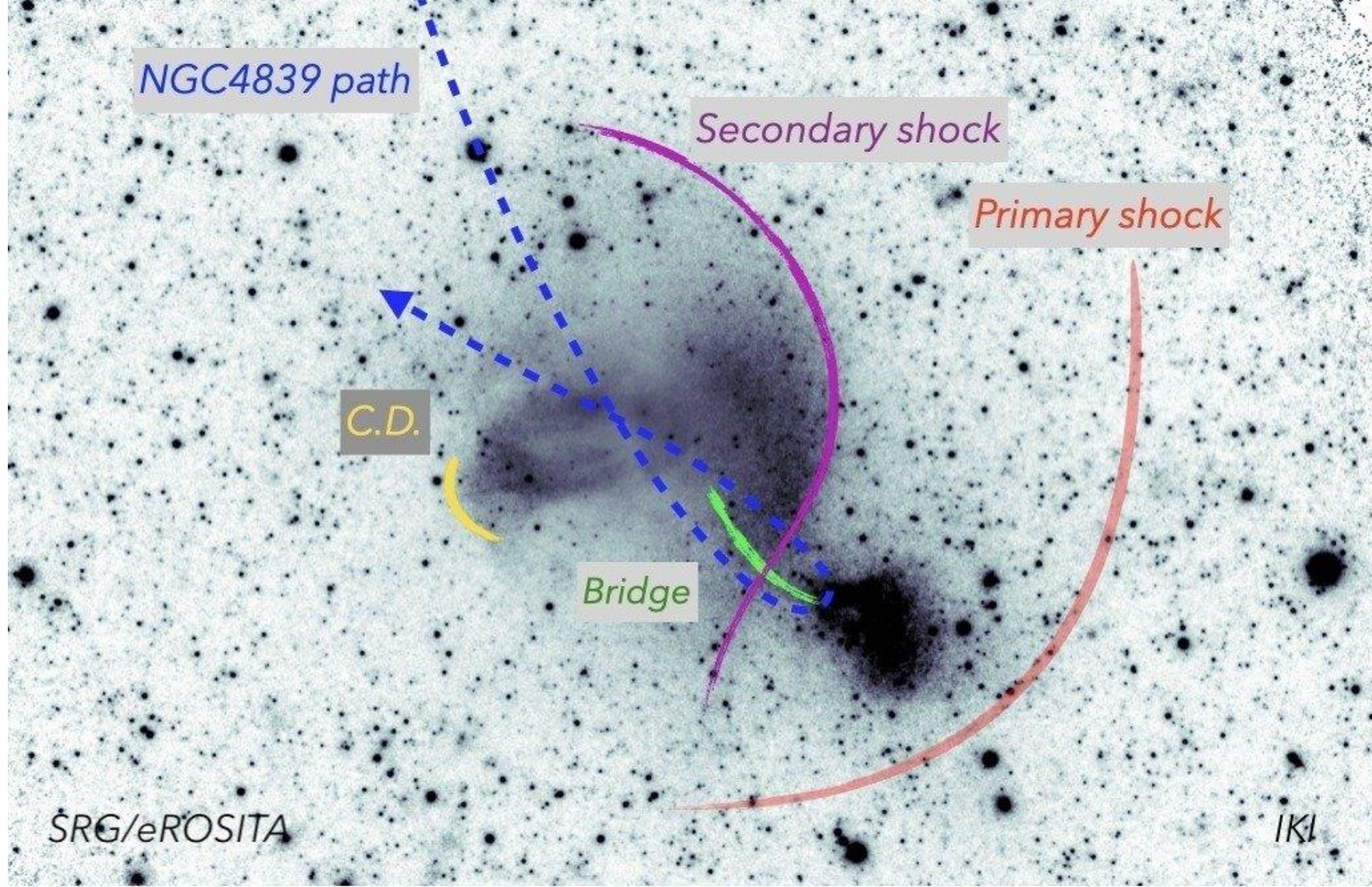
Number of X-Ray clusters
($0 < l < 180$):

	$\geq 5\sigma$	$> 4\sigma$
Sky survey 1	7,500	10,500
surveys 1+2	14,000	19,500
surveys 1+2+3	18,800	25,500
survey 1+2+3+4	23,200	31,500



eRosita after 8 sky surveys should discover majority of the massive clusters of galaxies in the »observable part of the Universe. : 50,000 clusters with mass $M > 2 \times 10^{14} M_{\odot}$ on the half of the sky or $\sim 100,000$ clusters of galaxies with many nearby group of galaxies)

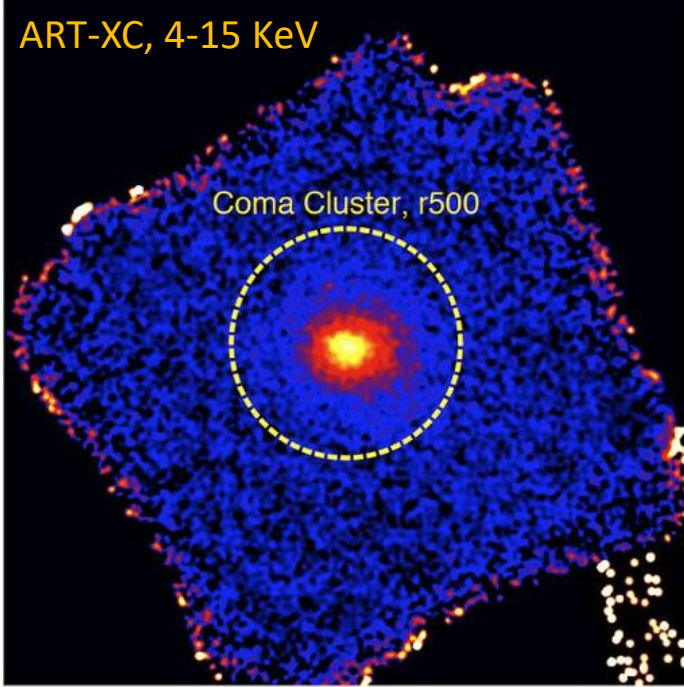
Marat Gilfanov and catalog group



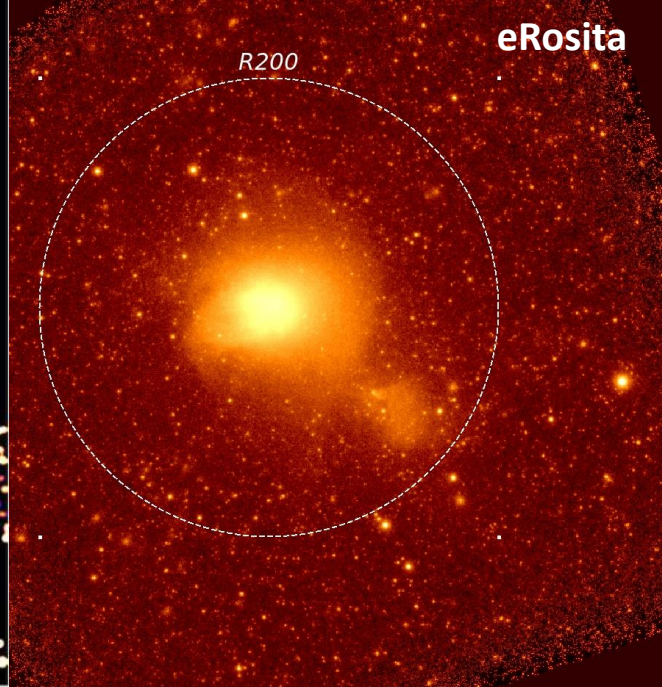
Churazov et al, A&A, 2021

Coma cluster, flattened image and suggested trajectory of the group NGC 4839

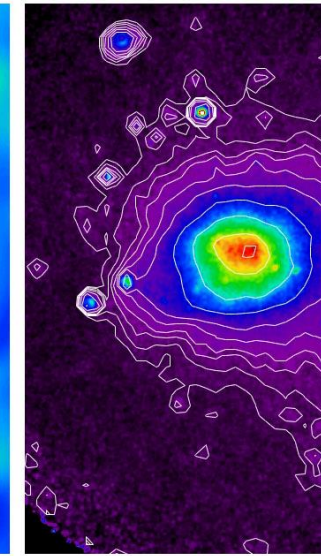
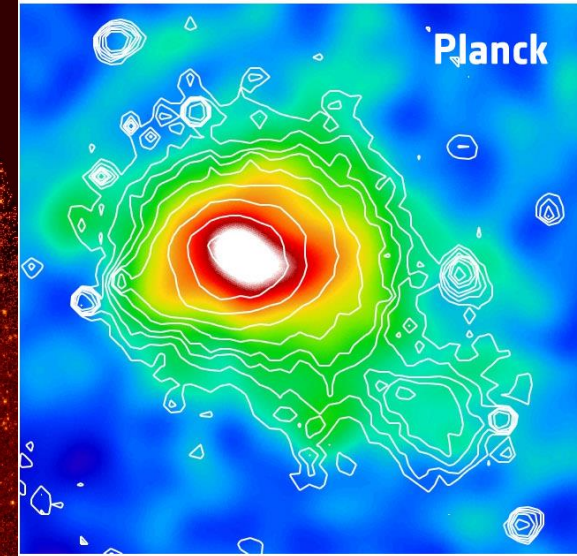
ART-XC, 4-15 KeV



eRosita



Planck

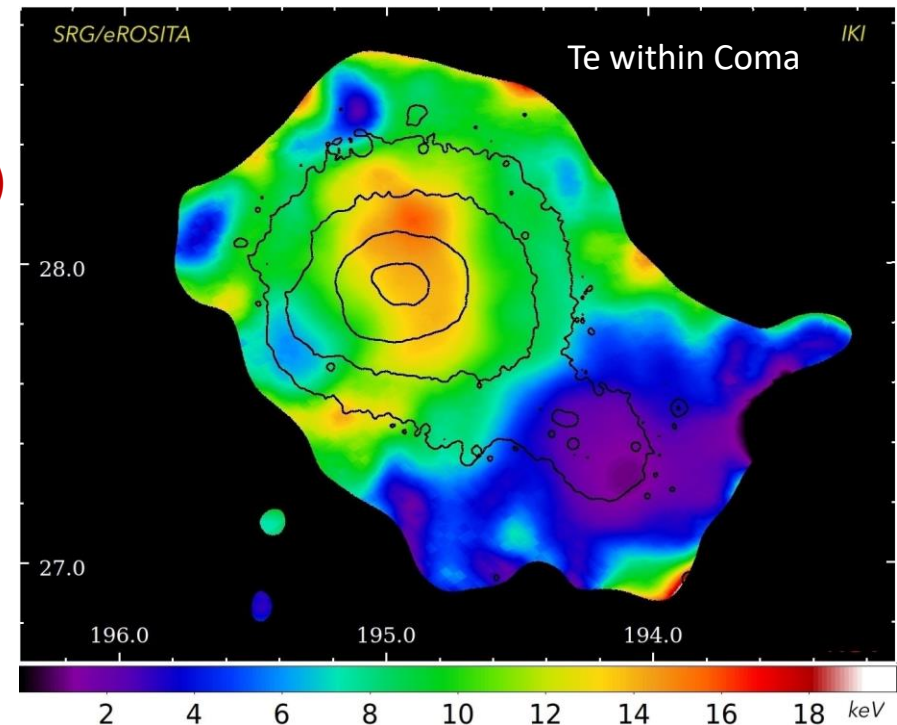


Excellent SZ-image

Ratio of “y” parameter on the Planck SZ map to the X-Ray surface brightness (eRosita, 0.4-2 KeV) opens the possibility to look for the distribution of Te over the whole Coma (under assumption of the gas density distribution)

$$k\tilde{T} \approx \frac{y}{I_X} \frac{m_e c^2}{\sigma_T} \frac{1}{4\pi} \frac{1}{(1+z)^2} \epsilon_c \left(\frac{n_H}{n_e} \right) \frac{\int n_e^2 dz}{\int n_e dz}$$

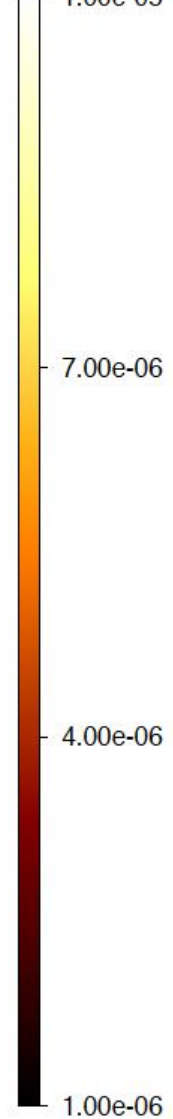
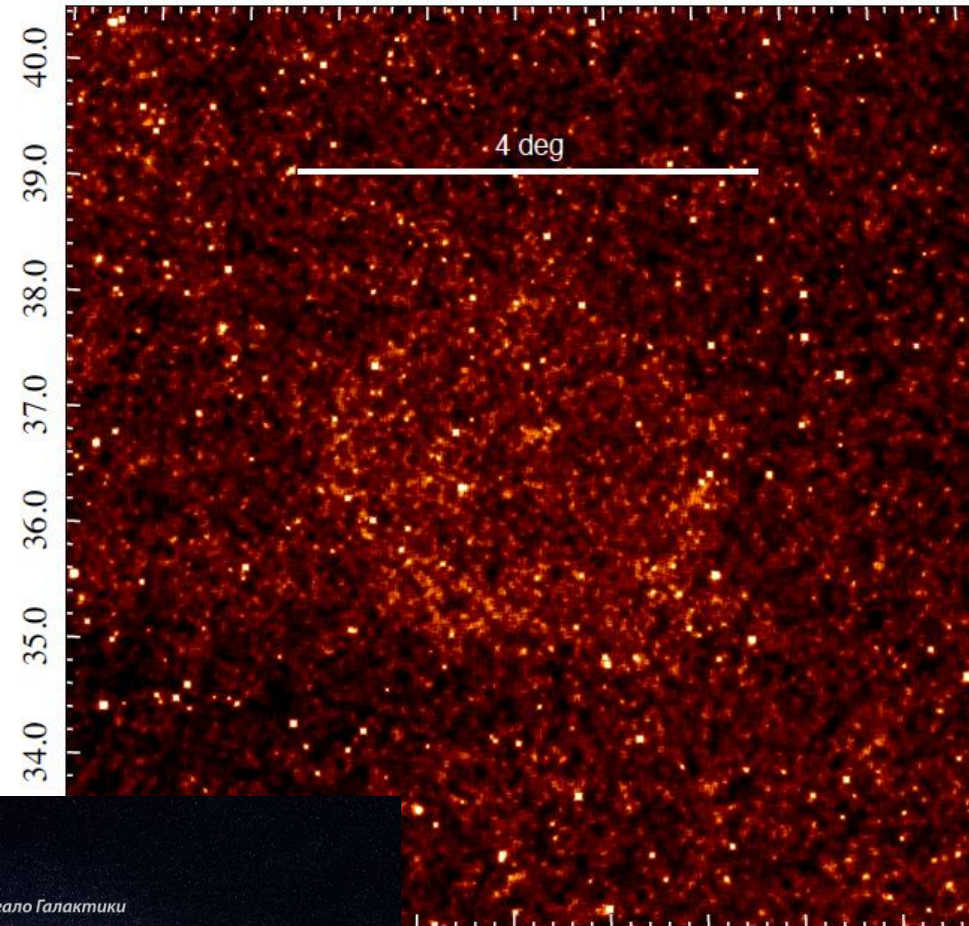
Churazov et al, A&A, 2021



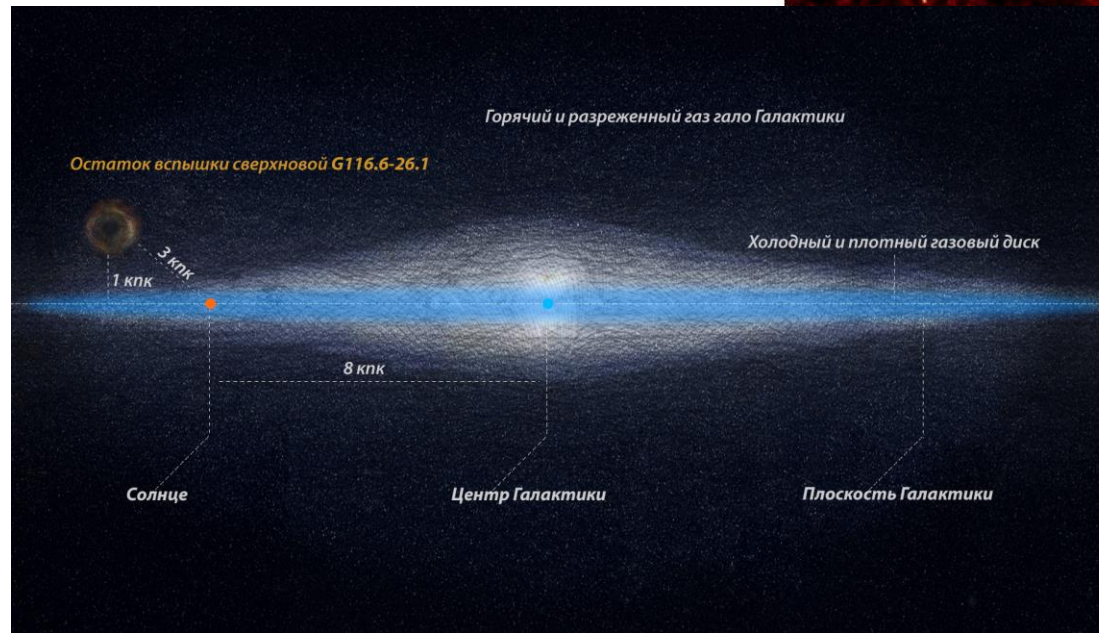
Very large remnant of the
thermonuclear supernova
explosion in the halo of our
Galaxy (1.3 kpc above plane
of the Galaxy, 3kpc from the Sun)

No traces in radio,
gamma-rays, optical
or infra-red
spectral bands

Declination



Right ascension



Churazov et al, MN RAS, 2021

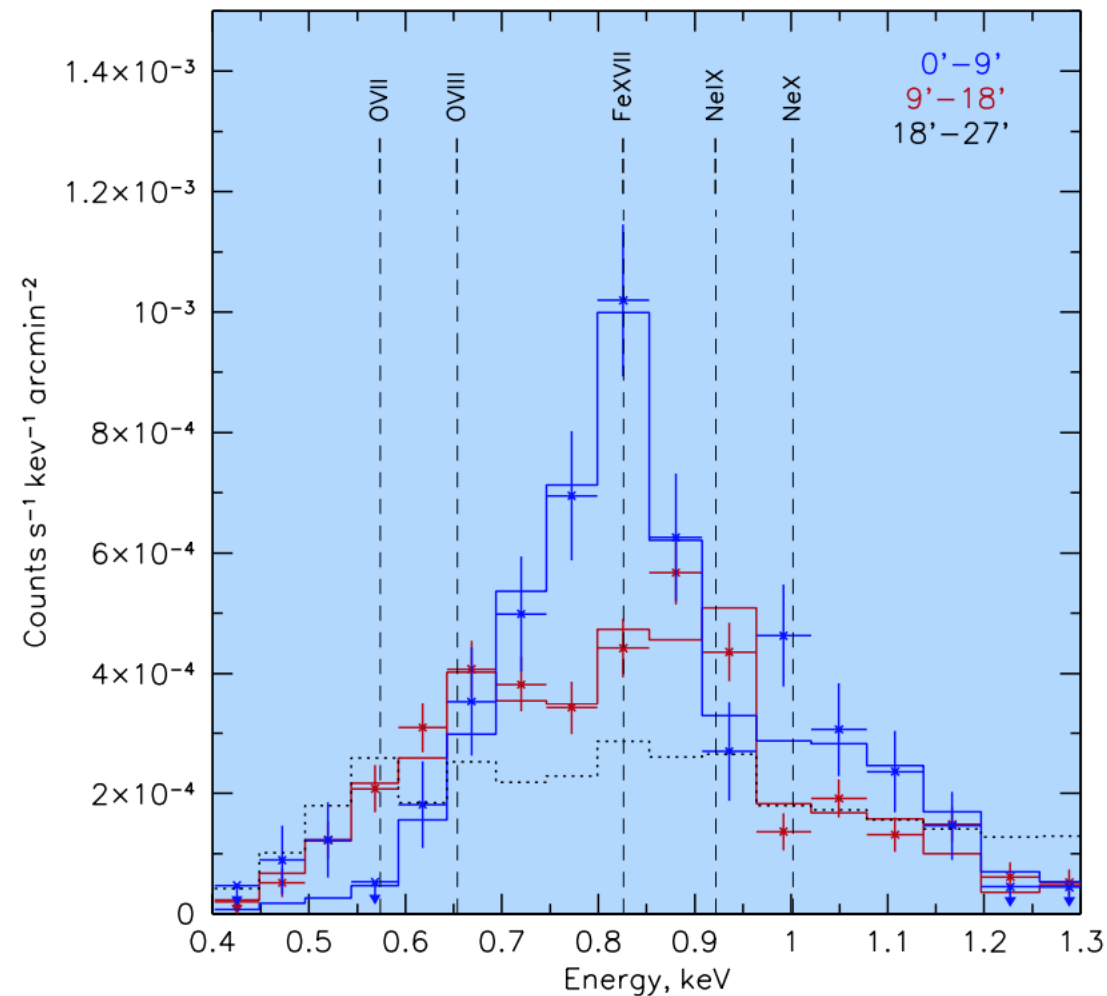
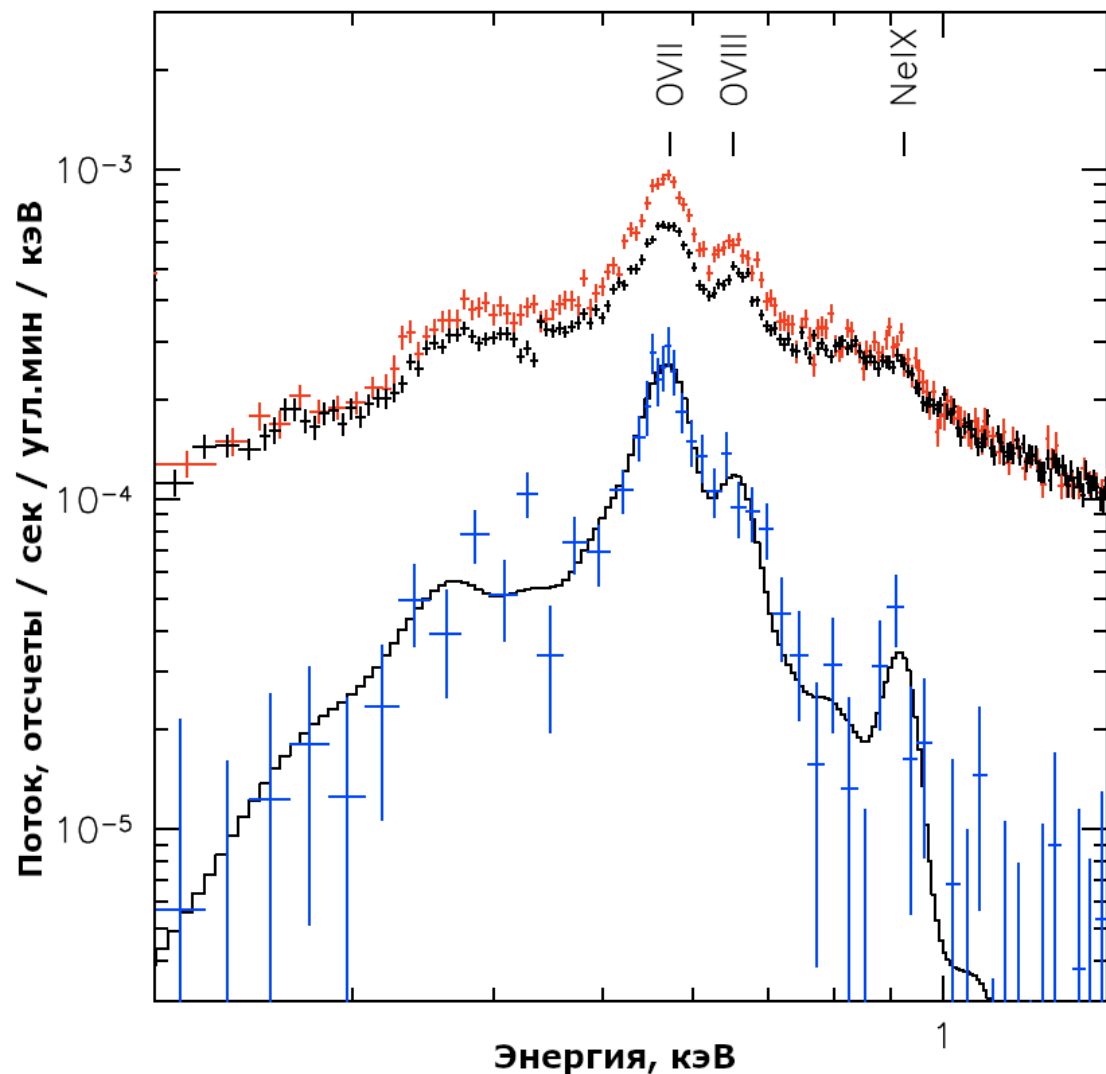
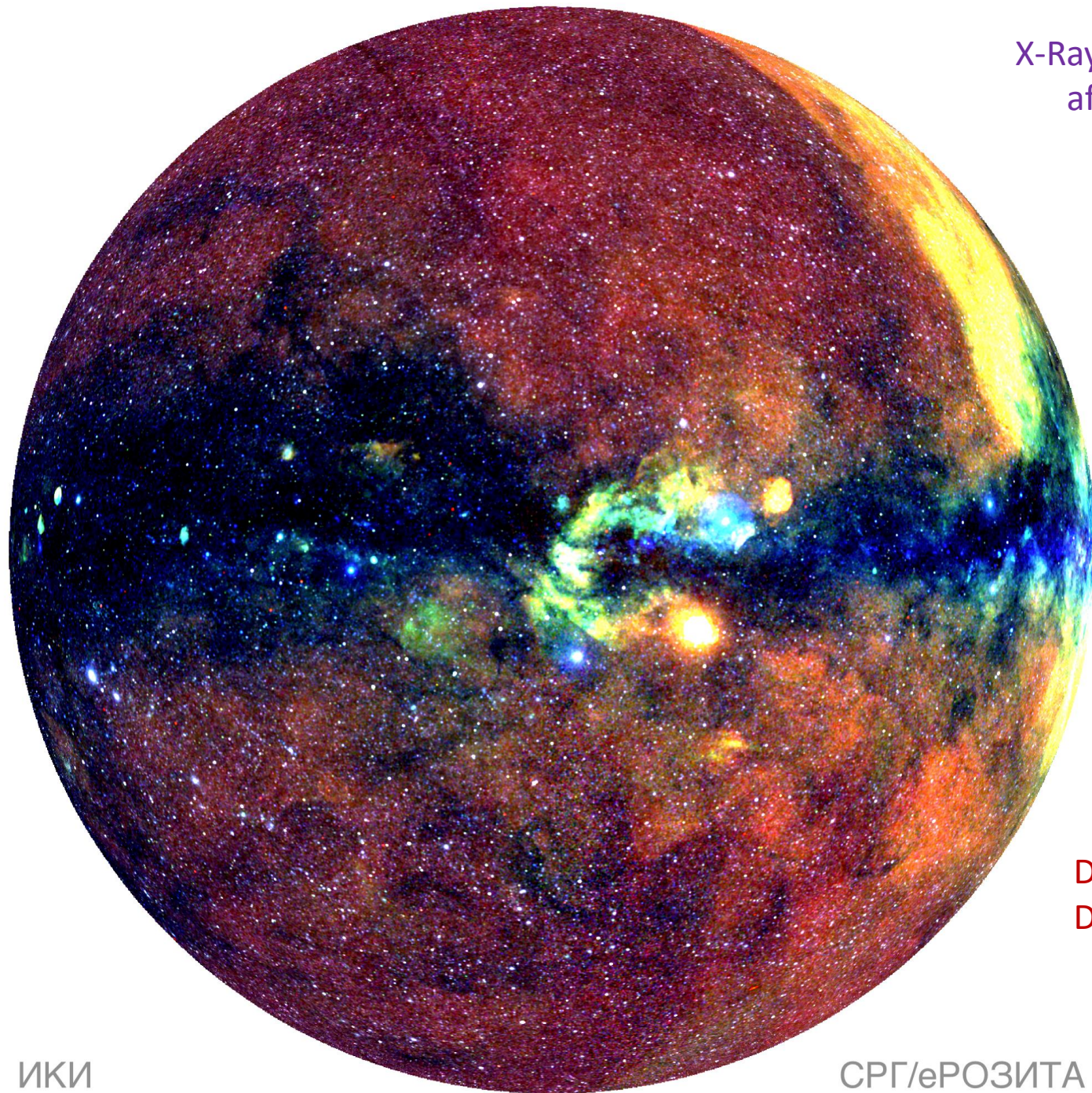


Figure 6. A zoom-in in the spectra (after astrophysical background subtraction) in the 0'-9' (blue) and 9'-18' (red) rings within the 0.4-1.3 keV band. The positions of the most prominent emission lines (O VII at 574 eV, O VIII at 654 eV, Fe XVII at 826 eV, Ne IX at 922 eV and Ne X at 1002 eV) are marked and labelled accordingly. The subtracted astrophysical background (and foreground) emission is shown by black dotted line.



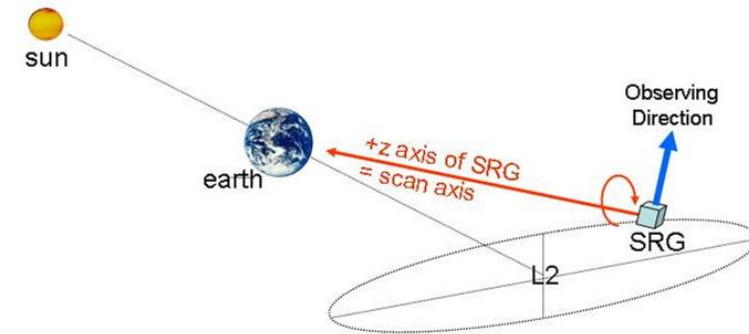
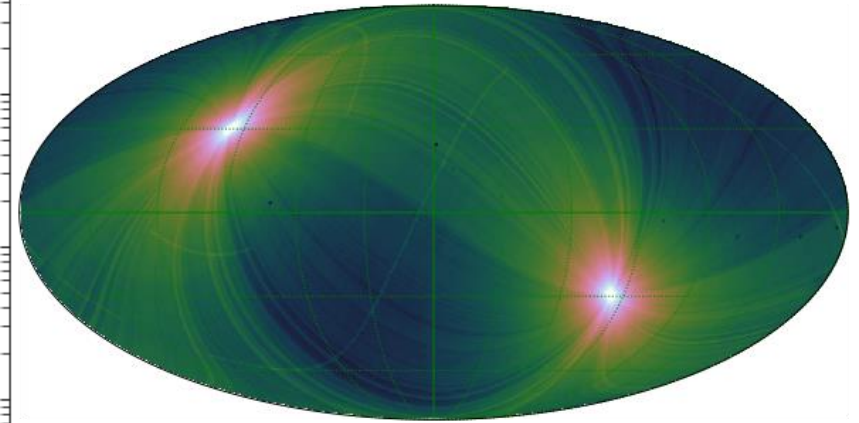
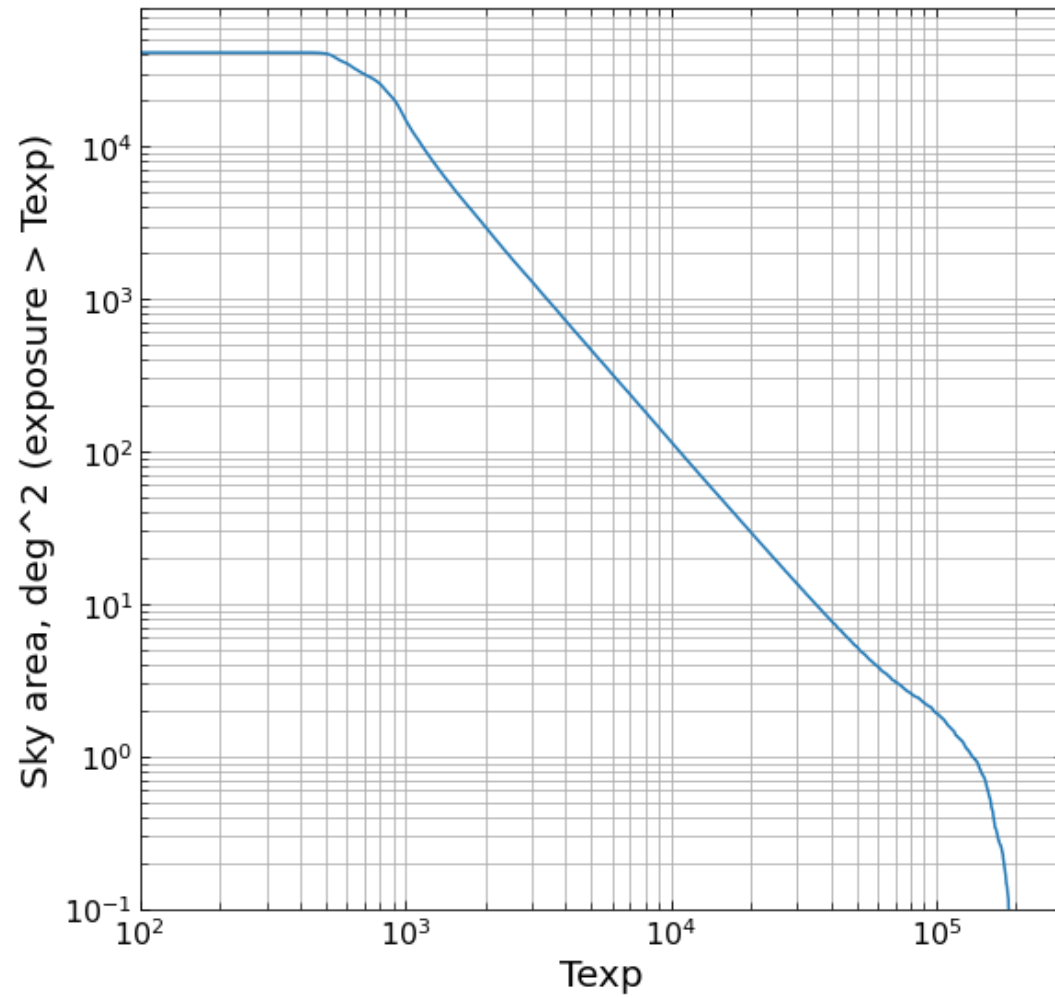
X-Ray map of one hemisphere
after two sky surveys

Amount of the sources
doubled: nearly
one million
on the one hemisphere

Strong variability:
tens of thousands
objects changed their
brightness during half
a year!

Discovery of the Tidal
Disruption Events (TDE)

Exposition map: 4 sky surveys: 12 Dec. 2019 –19 Dec. 2021



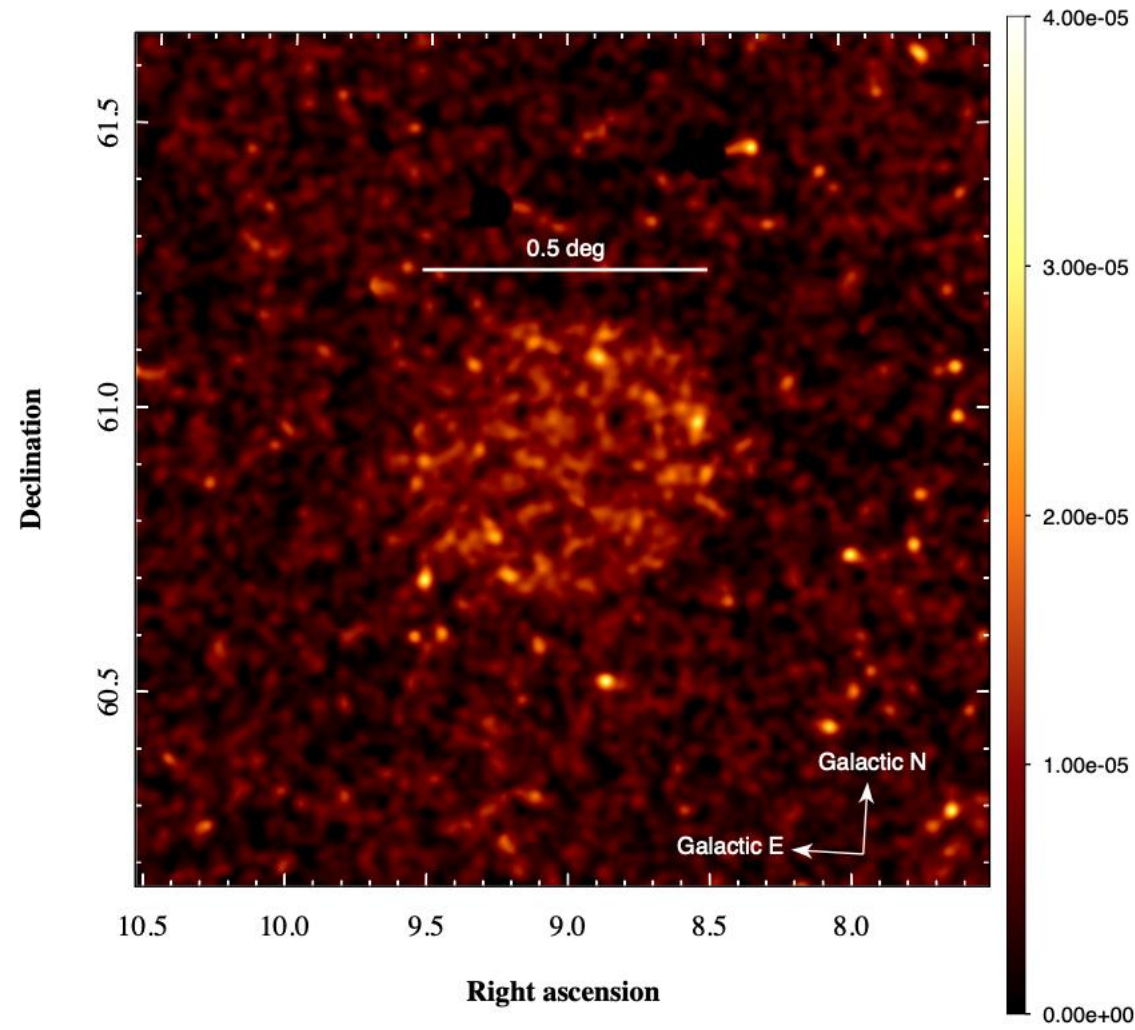
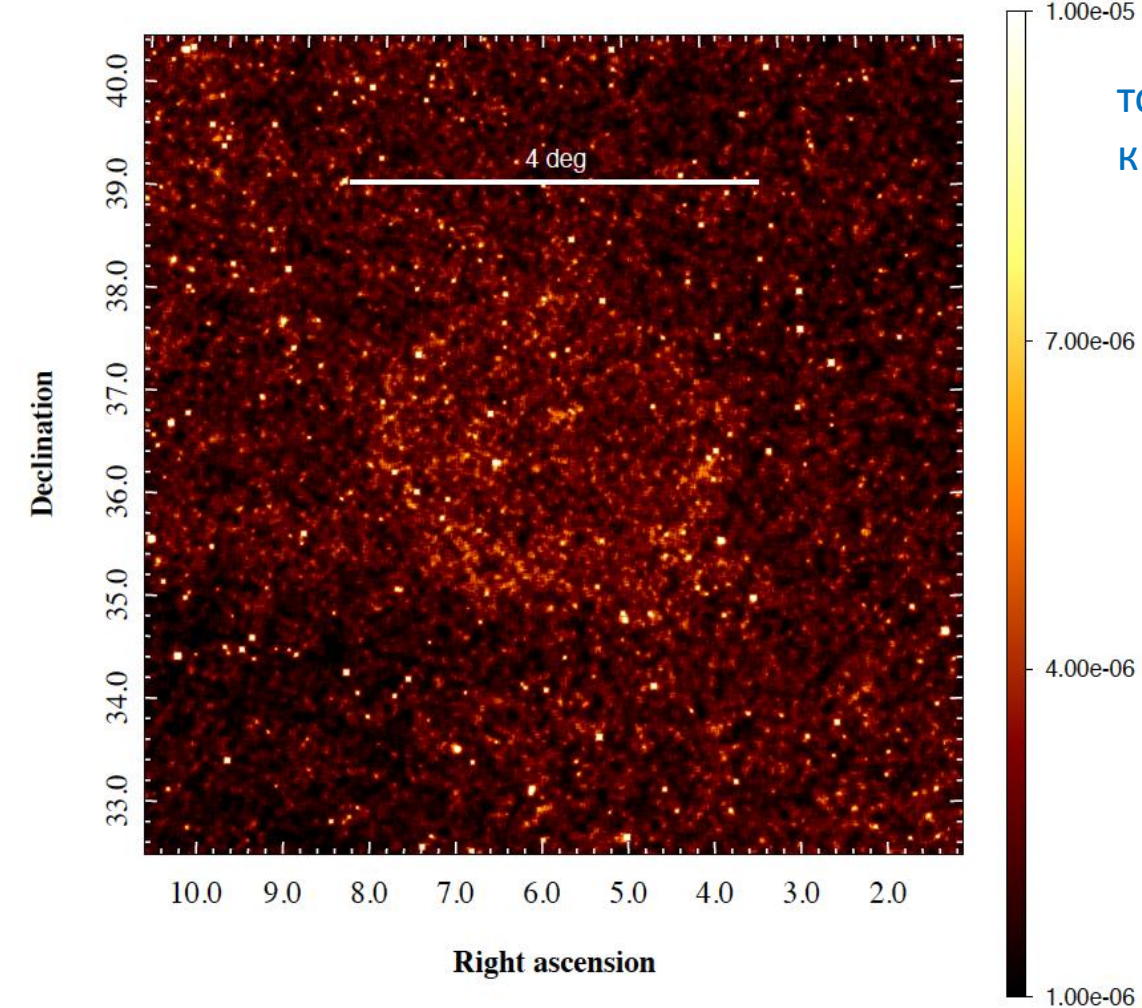
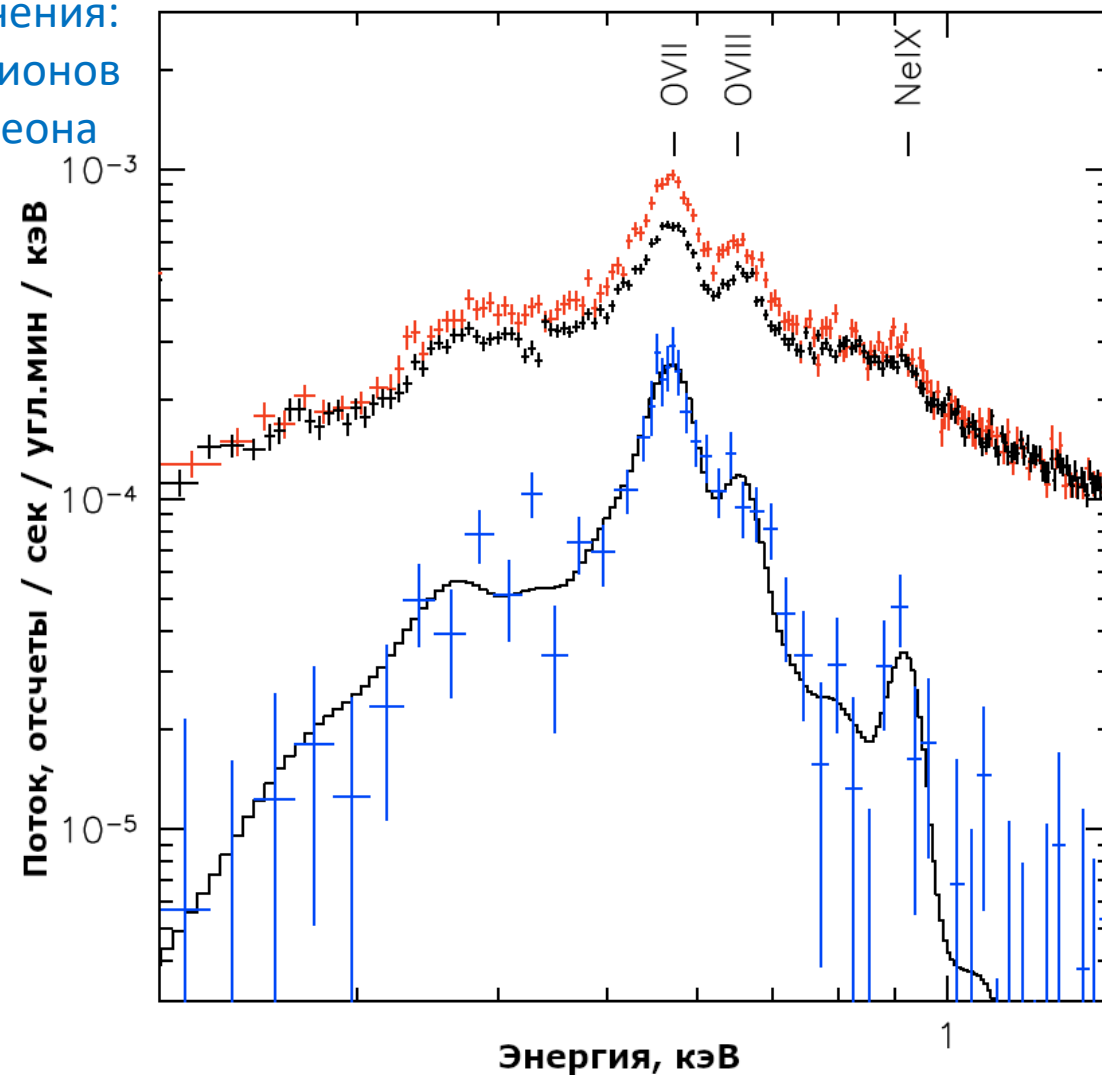


Figure 1. Detector background-subtracted vignetting-corrected image of X-ray (0.4-2.3 keV) surface brightness in the 1.5×1.5 deg field containing the newly discovered SNR candidate SRGe J003602.3+605421=G121.1-1.9 obtained by *SRG*/eROSITA after four surveys. The image was smoothed with a $\sigma=30''$ gaussian filter after masking point-like and mildly extended sources with the 0.5-2 keV flux above 3×10^{-14} erg s $^{-1}$ cm $^{-2}$. The white compass shows direction of the Galactic coordinates.



Спектр излучения:
только линии ионов
кислорода и неона



**Большой (~ 4 град) остаток взрыва термоядерной сверхновой
в гало Галактики** 1.3 кпк над плоскостью, 3 кпк от Солнца

Никаких следов в радио, гамма, оптике или ИК

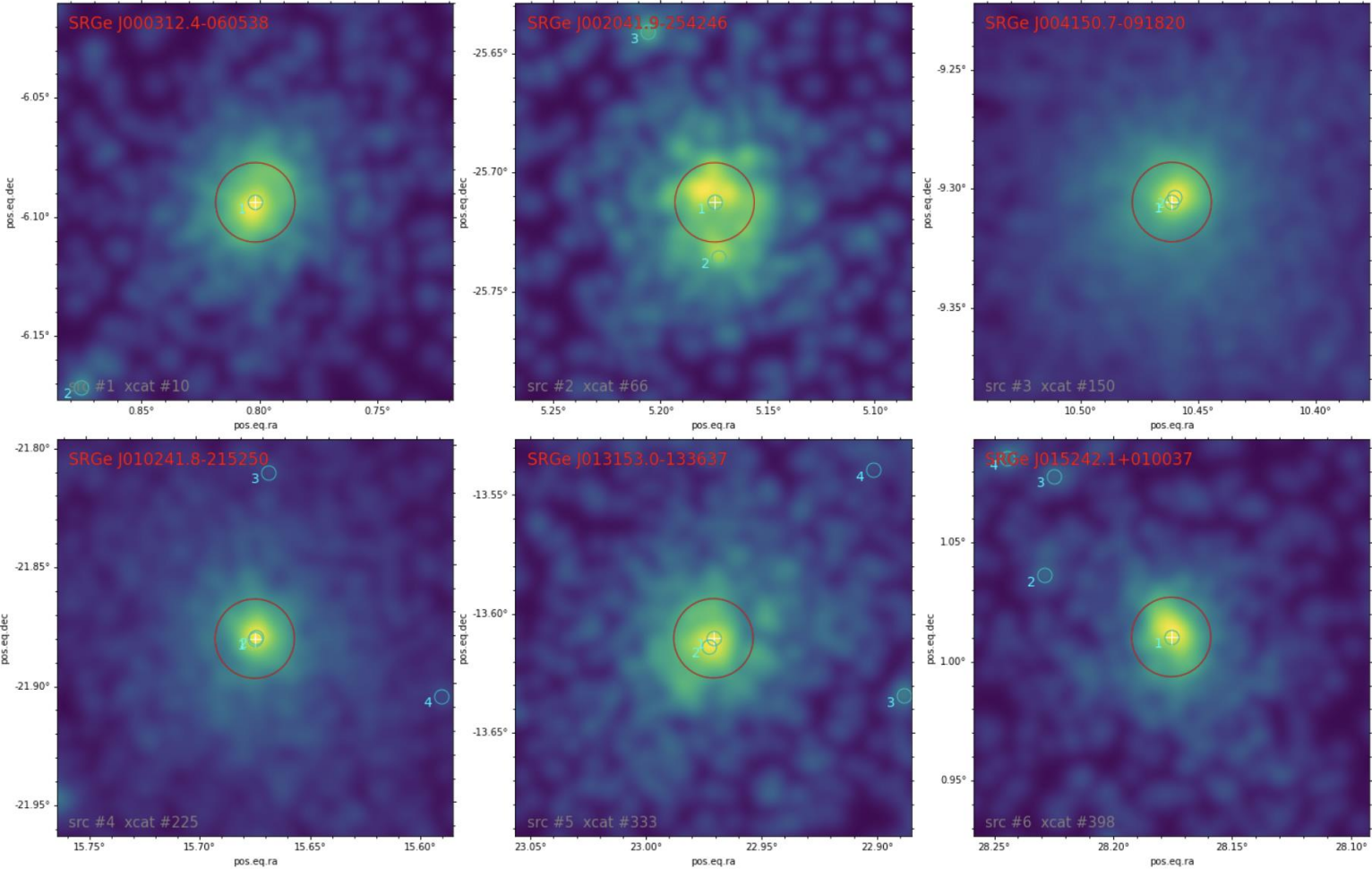
Пока только eРозита

**В статье И. Хабибуллина: о втором открытом
остатке сверхновой с сильными рентгеновскими
линиями ионов железа вблизи 826 эВ**

Скопления галактик, обнаруженные по их рентгеновскому излучению на половине неба (результаты eРозиты)

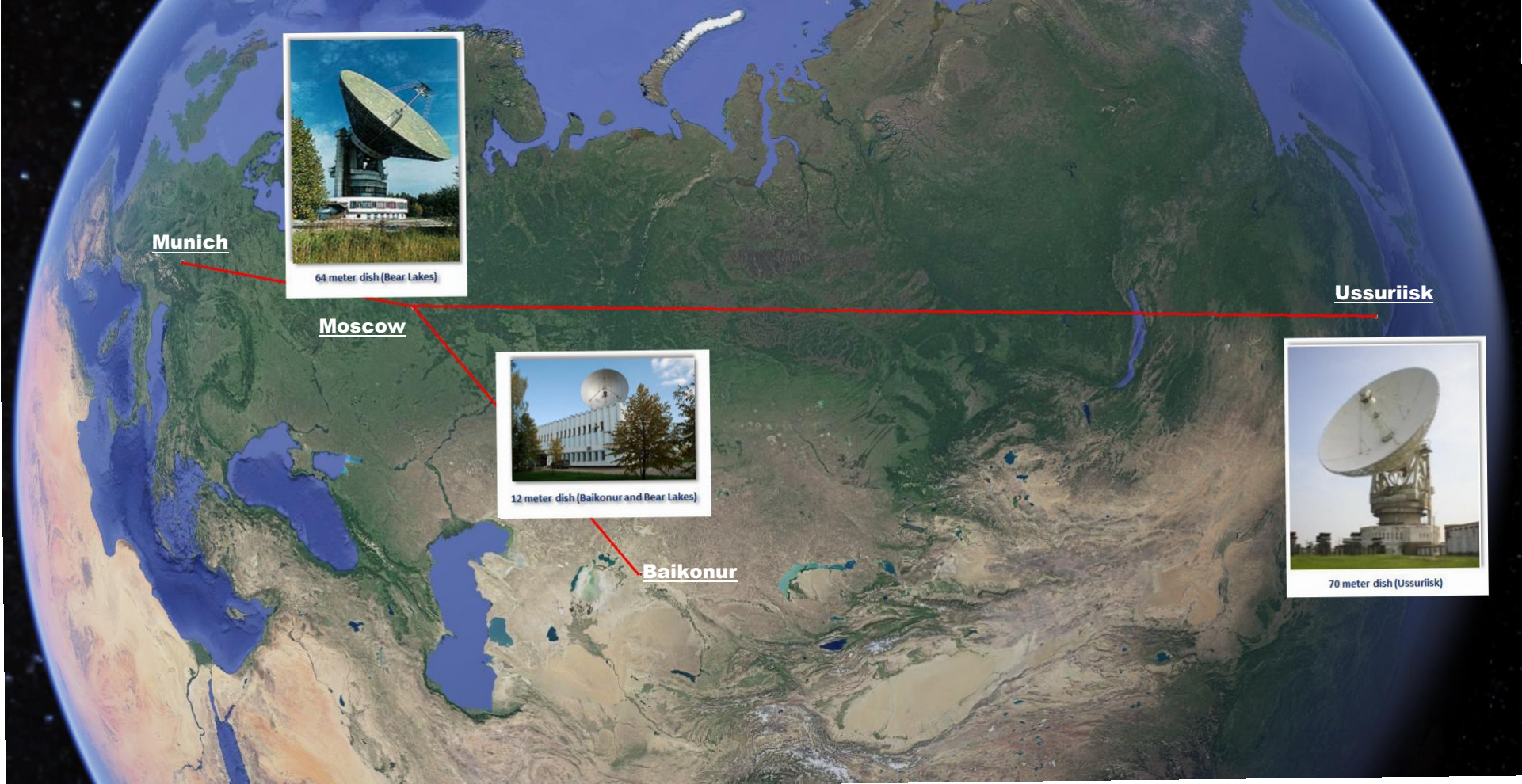
число обнаруженных скоплений
(0<l<180):

	$\geq 5\sigma$	$> 4\sigma$
обзор 1	7,500	10,500
обзор 1+2	14,000	19,500
обзор 1+2+3	18,800	25,500
обзор 1+2+3+4	23,200	31,500

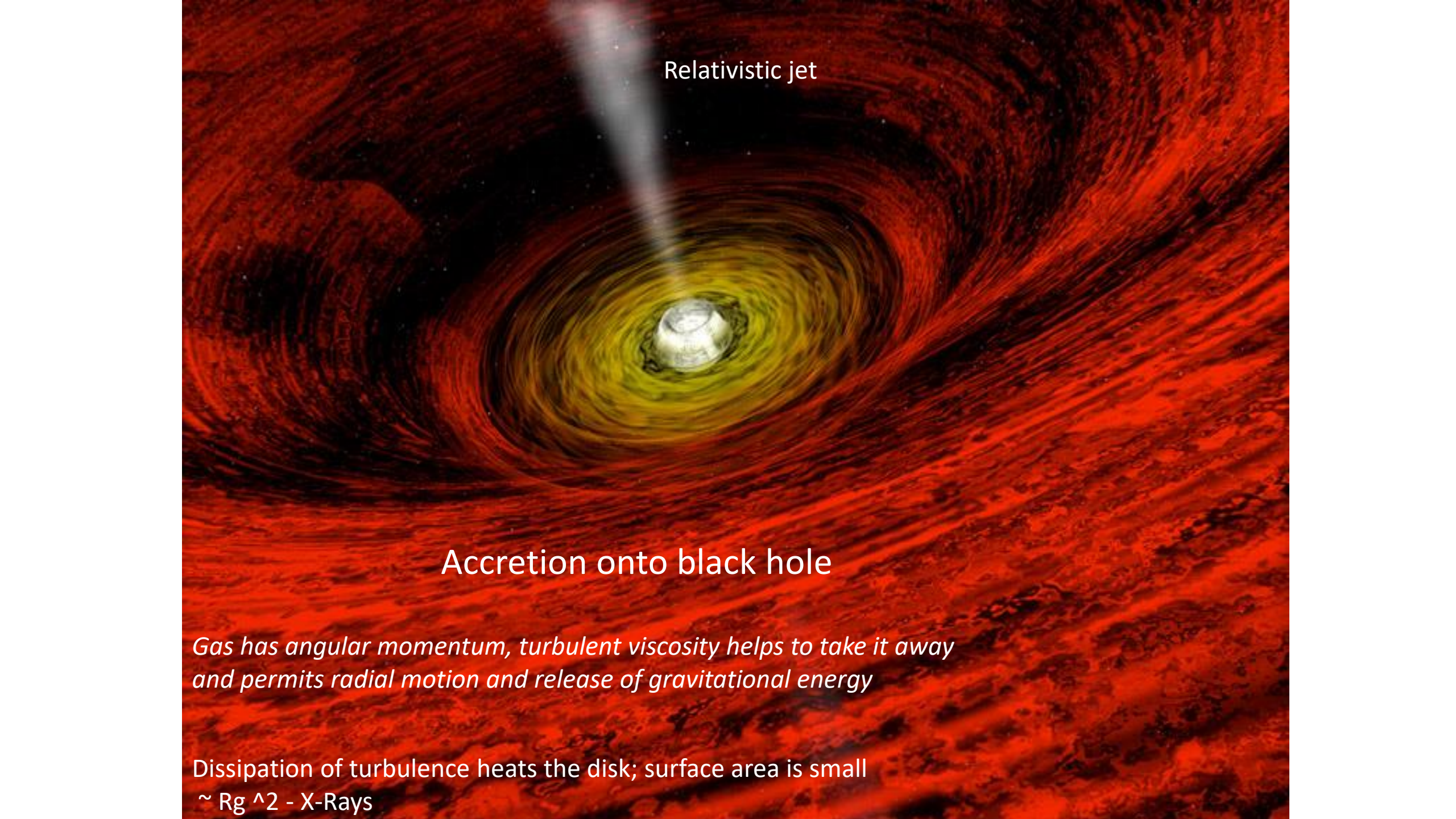


eРОЗИТА за 8 сканов должна обнаружить все массивные скопления галактик в наблюдаемой части Вселенной:
50,000 скоплений с вириальной массой $M > 2 \times 10^{14} M_{\odot}$
(всего ~100,000 сколений и групп галактик)

М.Р. Гильфанов и группа по каталогам



Antennae used by the Center of Deep Space Communications
for everyday work with SRG Observatory and its telescopes

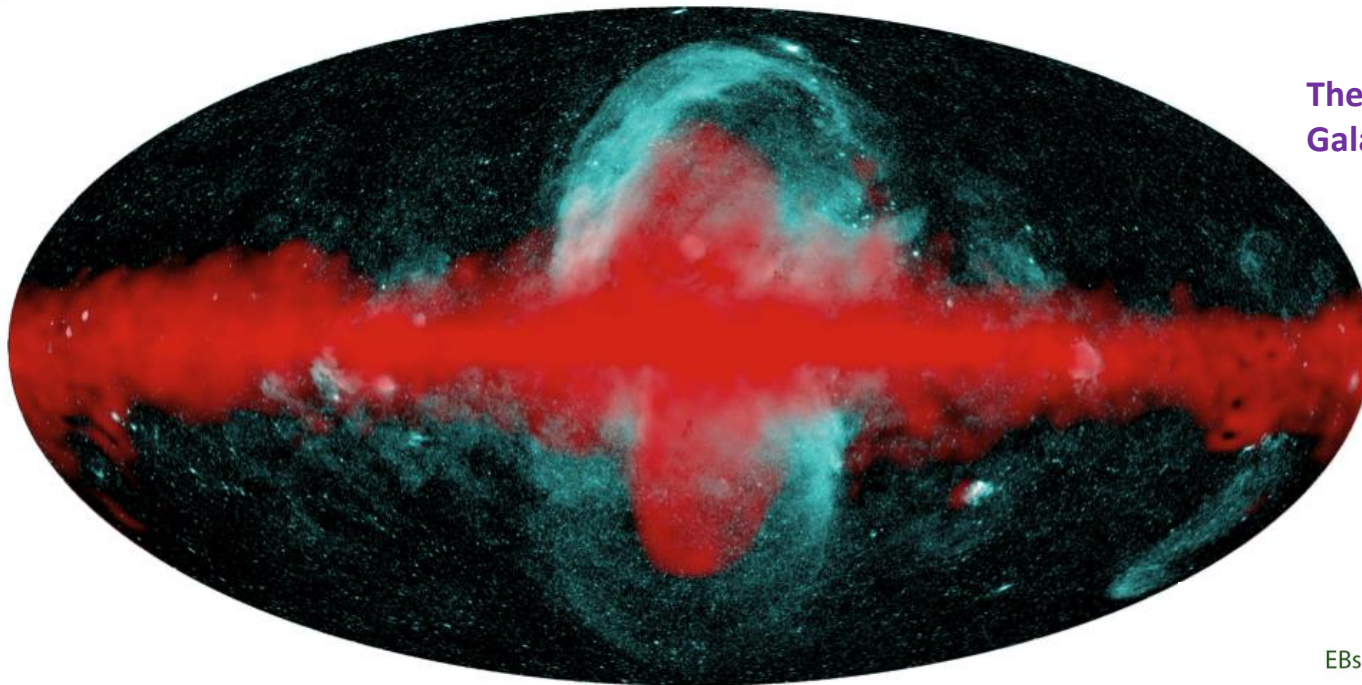


Relativistic jet

Accretion onto black hole

*Gas has angular momentum, turbulent viscosity helps to take it away
and permits radial motion and release of gravitational energy*

Dissipation of turbulence heats the disk; surface area is small
 $\sim Rg^2$ - X-Rays



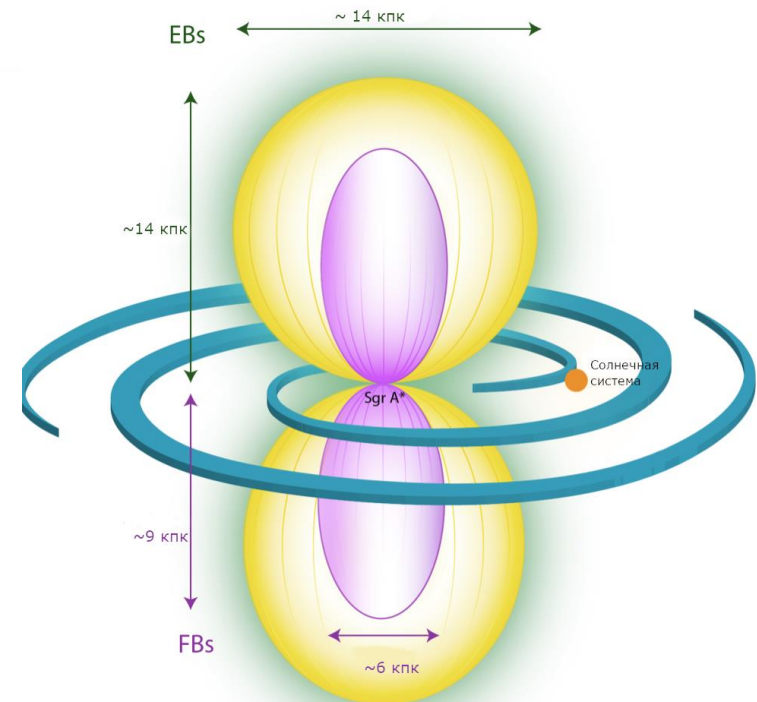
«eRosita bubbles»

The activity of the central part of our Galaxy tens of millions of years ago

Overlay of maps of our Galaxy obtained by telescopes **SRG / eROSITA** and **Fermi (NASA)**. Diffuse X-ray radiation recorded by the SRG / eROSITA telescope (0.6–1 keV, indicated by shades of blue), surrounds a region of harder radiation (GeV's), indicated in red, which is called “Fermi bubbles”.

Images from the article P. Predehl, R. Sunyaev, et al., Nature, 2020

sketch



Every day the plane of the scan shifts slowly to one degree, following the Sun and leaving one degree wide strip on the sky map



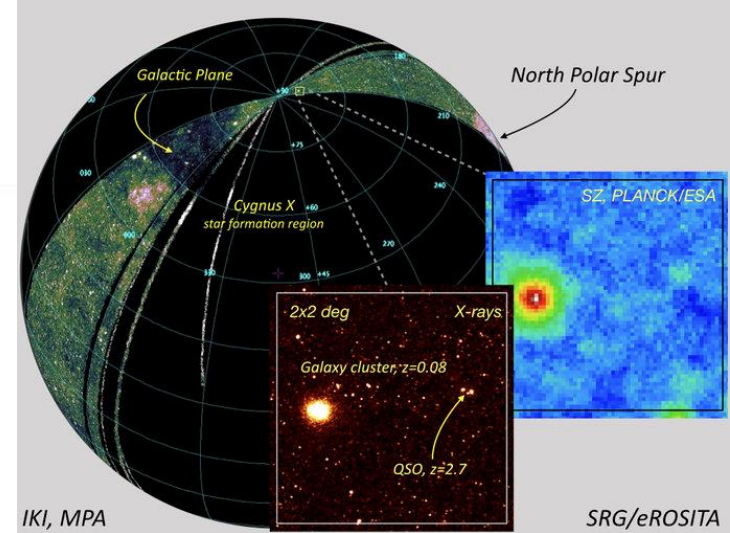
One revolution around direction toward Sun every 4 hours.



*+z axis of SRG
= scan axis*

Observing
Direction

SRG



Every source on the sky is observed 6 times per day once in a half year (*This permits to look for the short time variability*)

Lavochkin industry near Moscow
one of the clean rooms

ART-XC (350 kg)
Russia

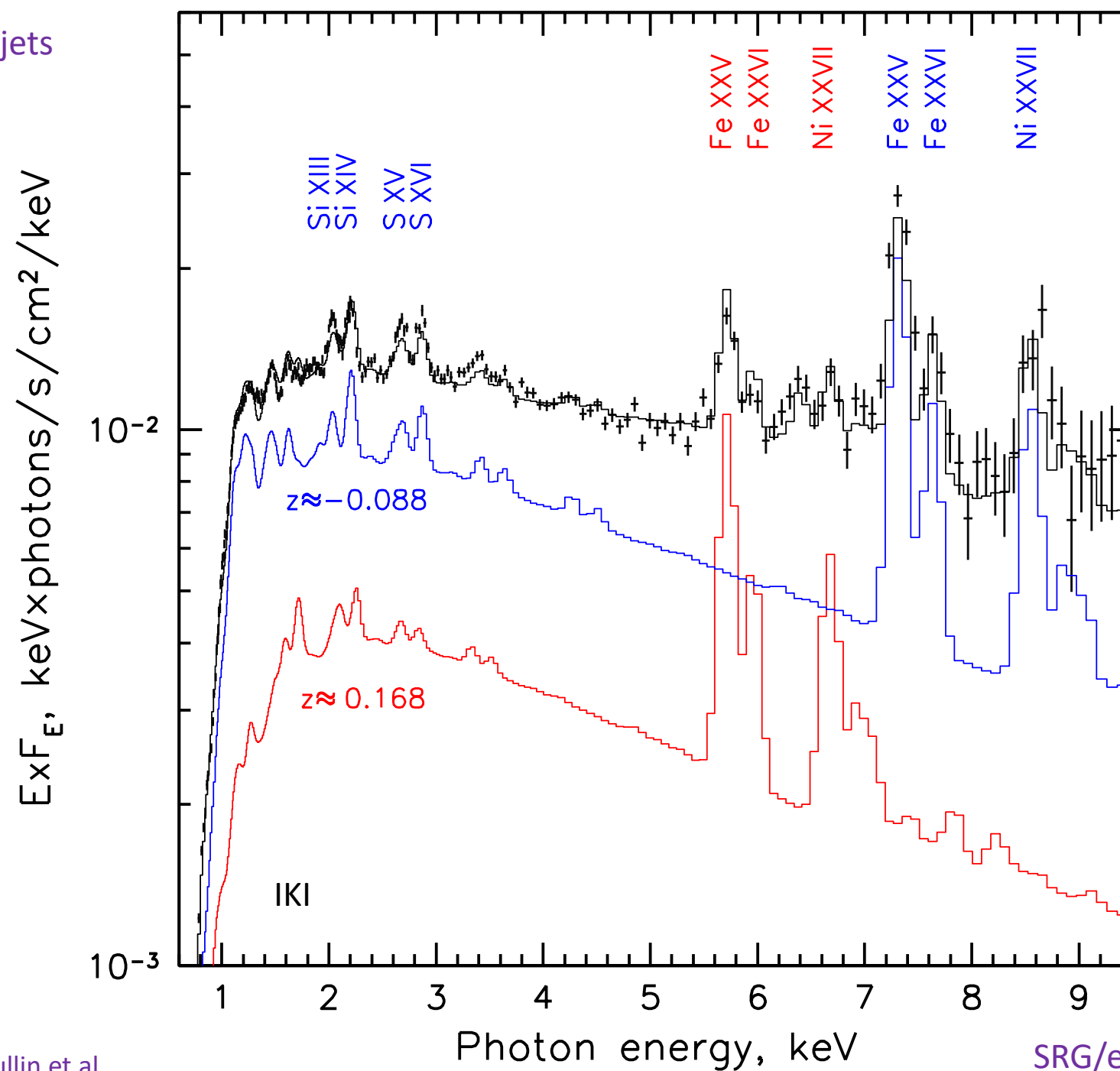
Navigator
Platform

SRG Observatory
total mass 2700 kg

eRosita (808 kg)
3.5m * 1.9m
Germany

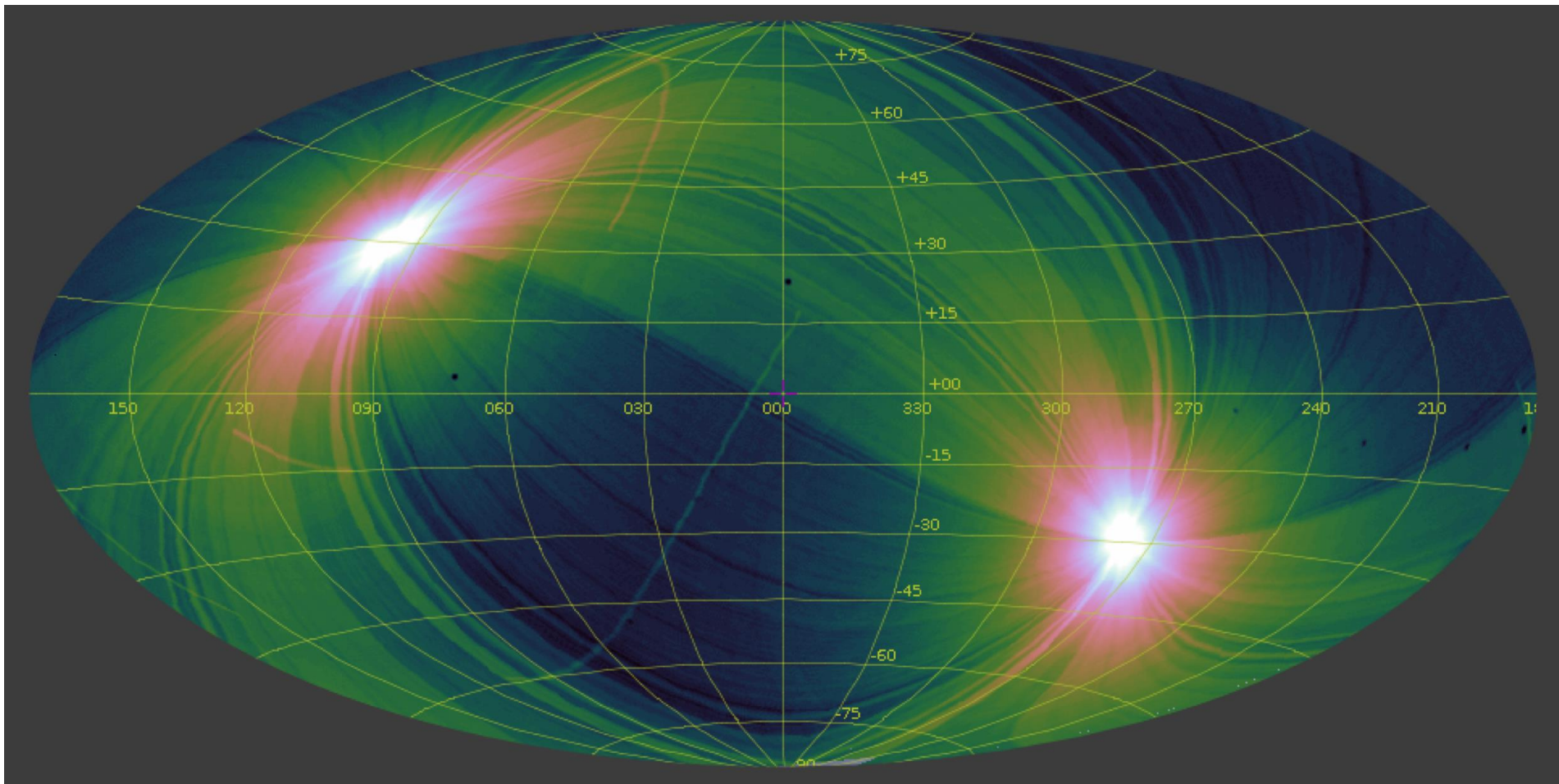


SS 433 jets



Khabibullin et al

SRG/eROSITA



Exposition map (galactic coordinates)

all individual scans intersect
in Ecliptic Poles

